

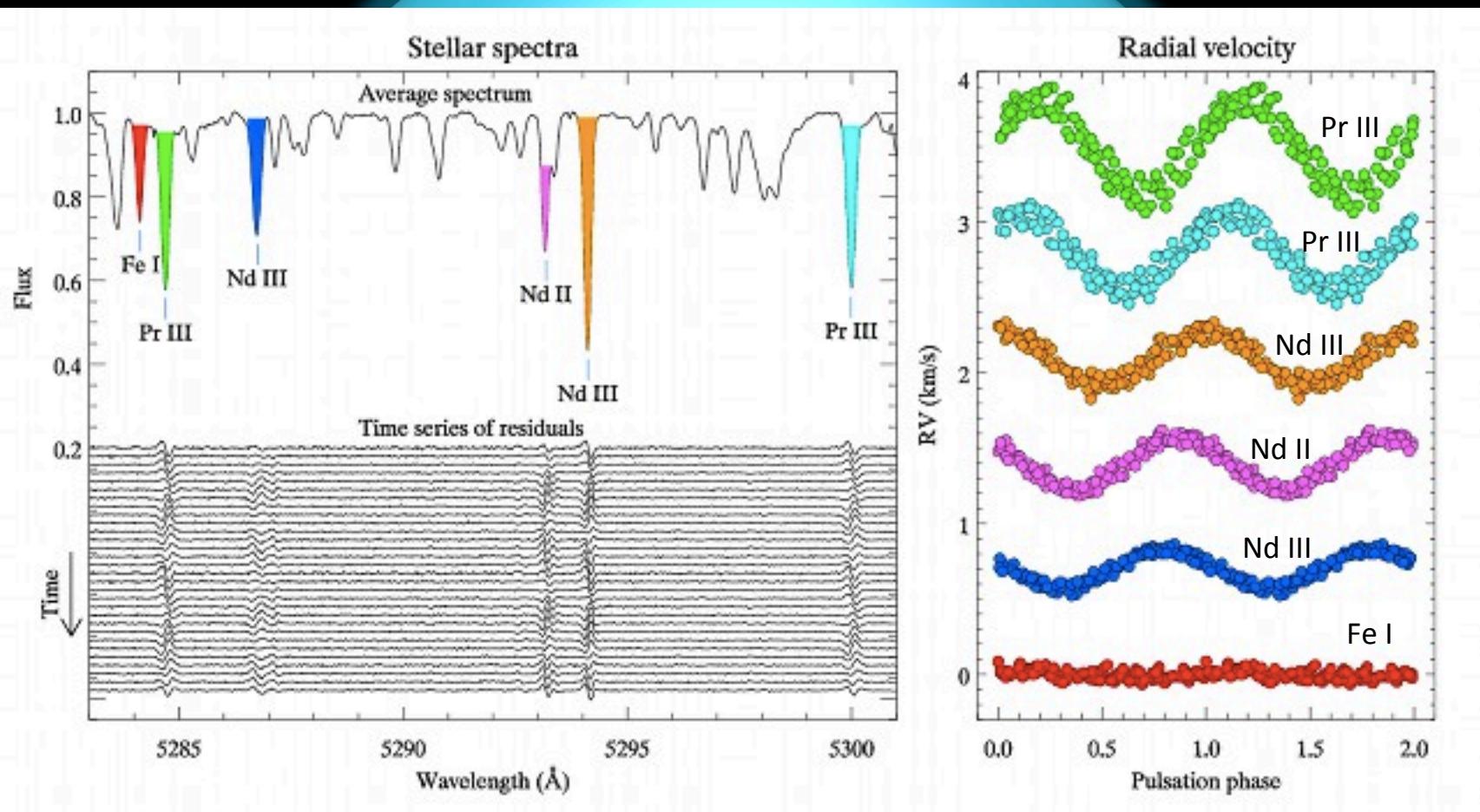
A theoretical tool for the study of radial velocities in the atmospheres of roAp stars

Paola Quirral-Manosalva and Margarida S. Cunha

Faculdade de Ciências – Universidade do Porto

IA-Porto Instituto de Astrofísica e Ciências do Espaço





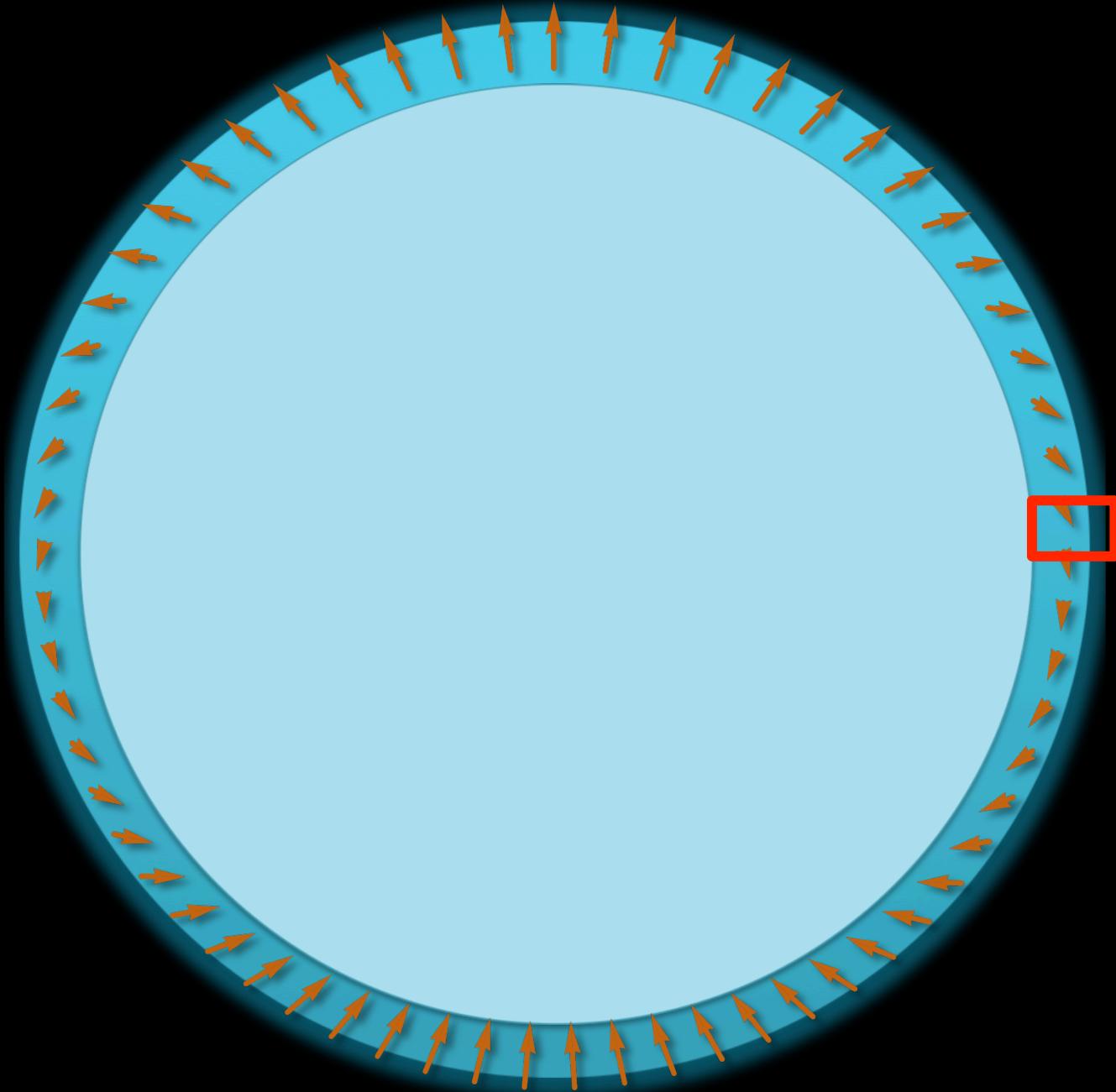
Oscillation Model:

$$\frac{D\rho}{Dt} + \rho \nabla \cdot \vec{v} = 0$$

$$\rho \frac{D\vec{v}}{Dt} = -\nabla p + \rho \vec{g} + \vec{j} \times \vec{B}$$

$$\frac{Dp}{Dt} = \frac{\gamma p}{\rho} \frac{D\rho}{Dt}$$

$$\frac{\partial \vec{B}}{\partial t} = \nabla \times (\vec{v} \times \vec{B})$$



Stellar model:

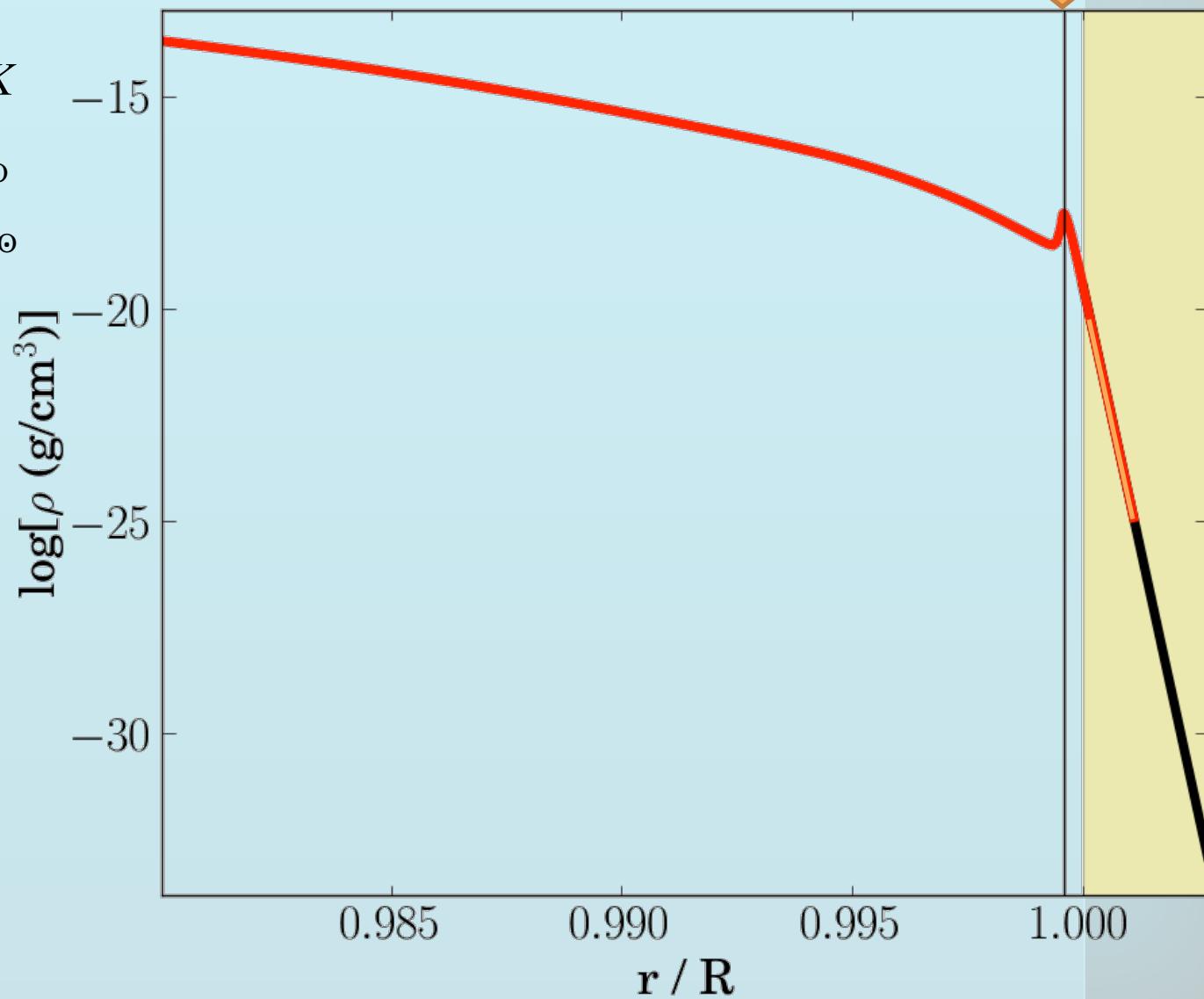
$T_{eff} = 8363 K$

$M = 1.8 M_\odot$

$R = 1.57 R_\odot$

THE PHOTOSHERE IS HERE

Ph

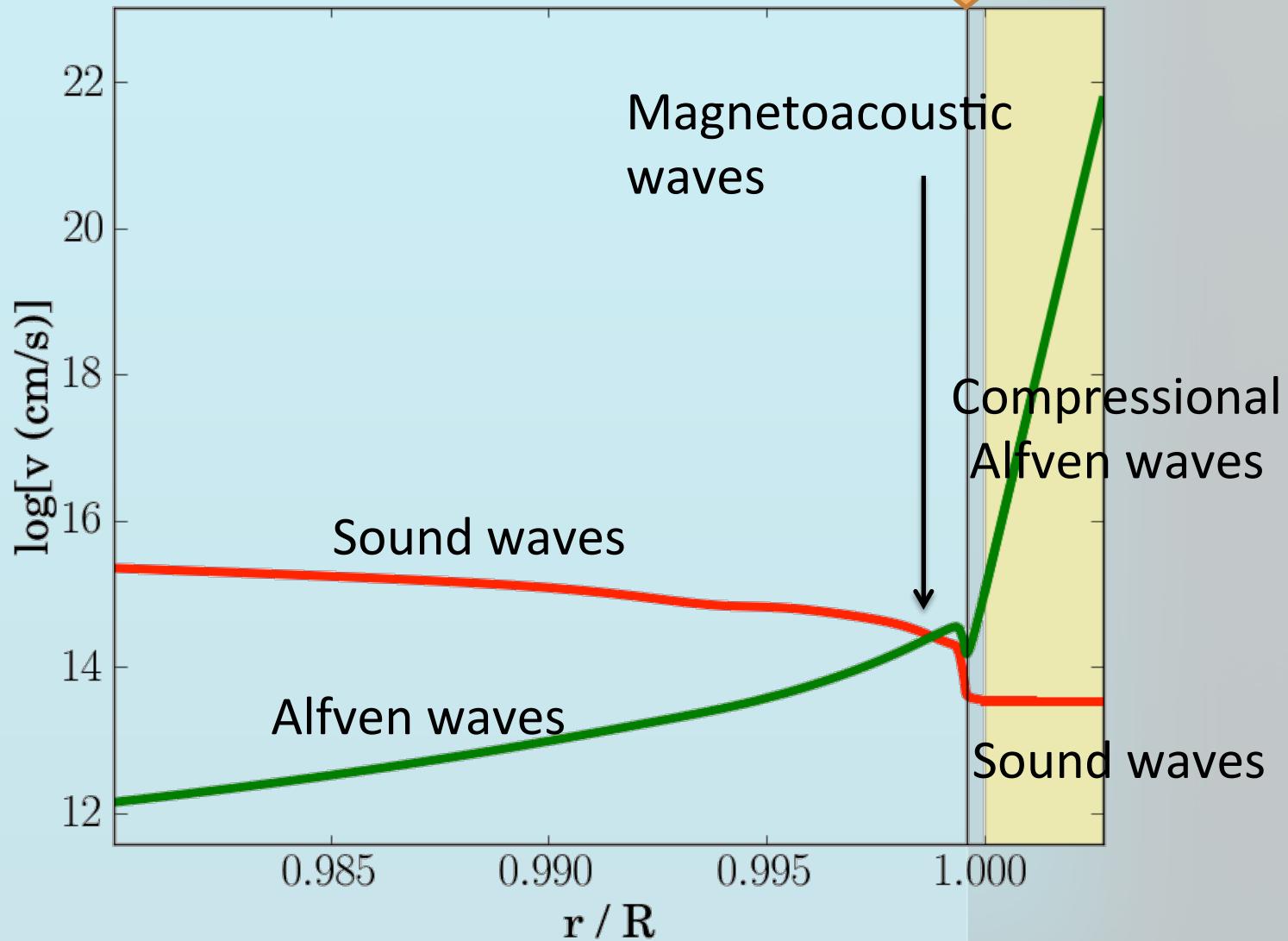


Sound velocity vs Alfvén velocity

in red

in green

Ph



Calculation of the oscillation solutions:

Numerical solution:

0.975

0.980

0.985

0.990

0.995

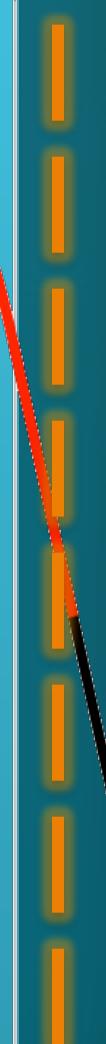
1.000

R

Cunha 2006

Sousa & Cunha 2011

Analytical solution:

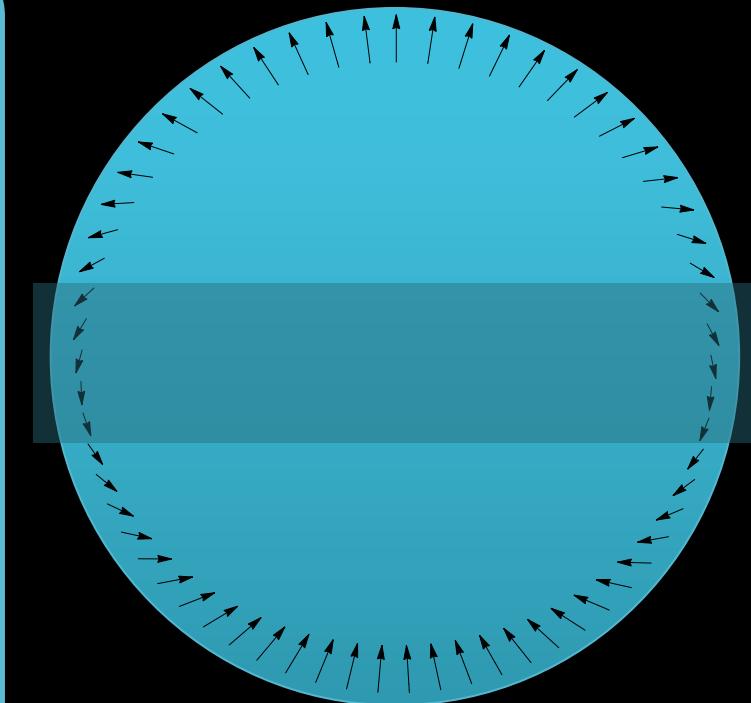


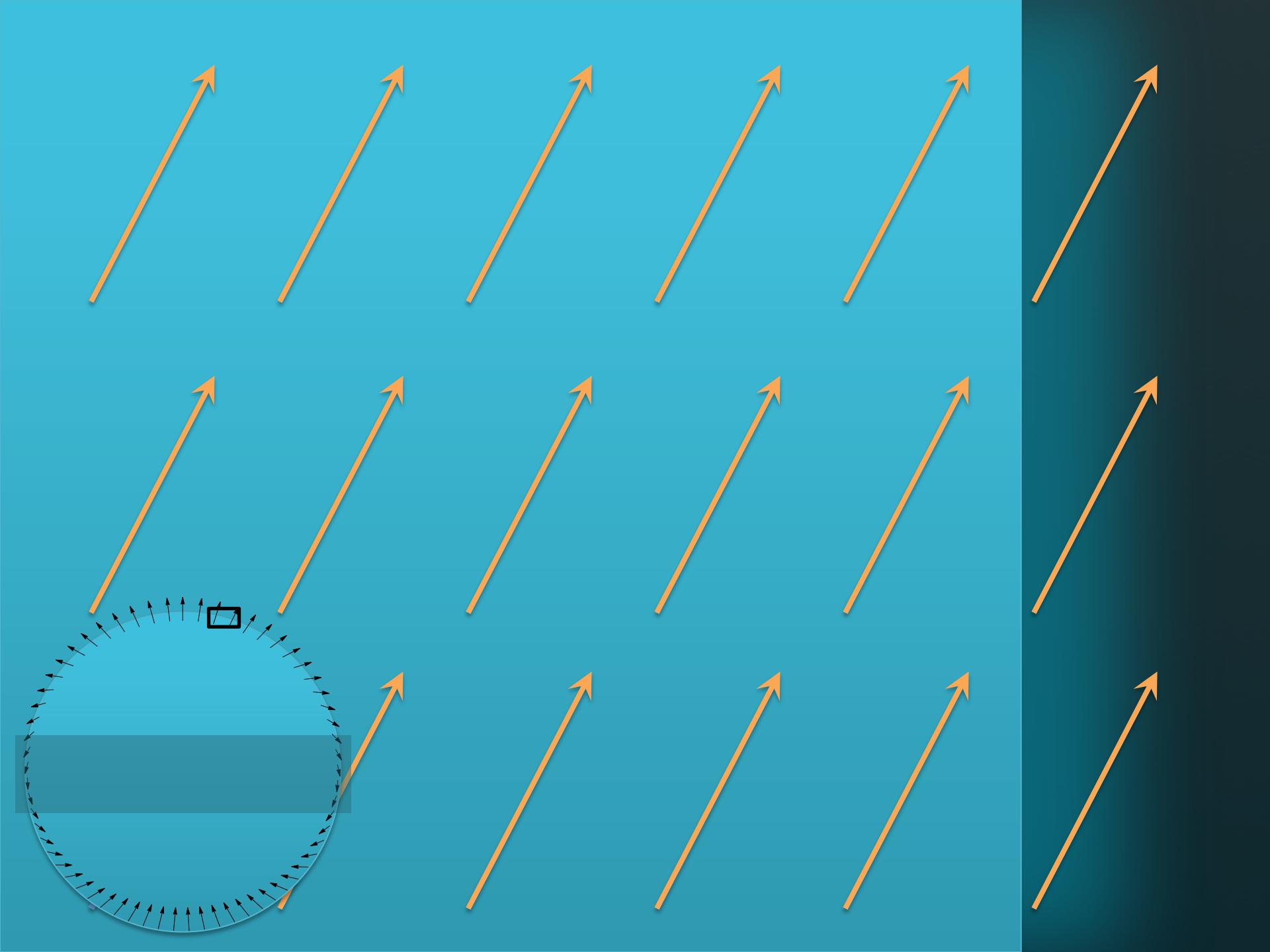
Analytical solutions:

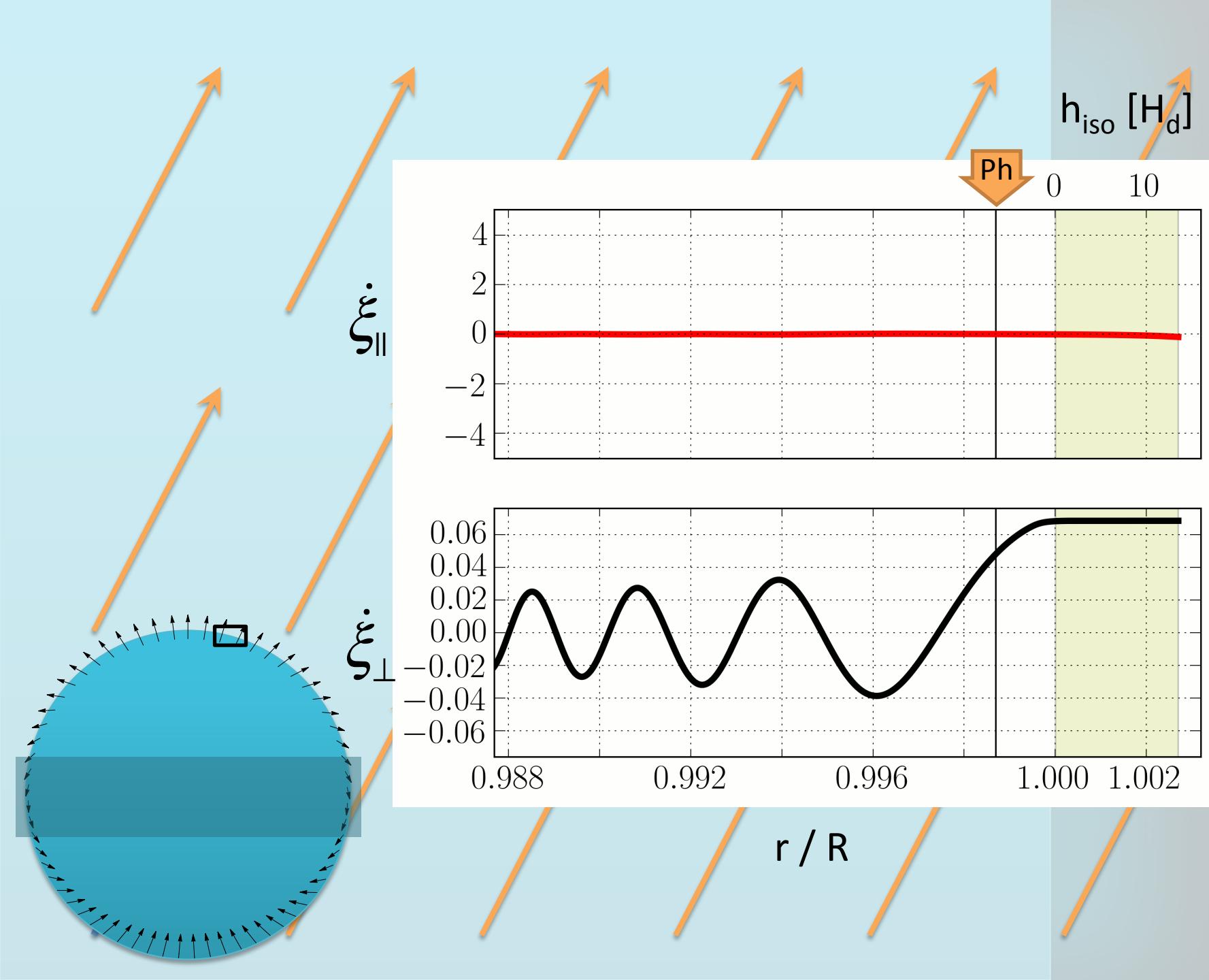
$$\xi_{\perp} = A_f J_0(\nu_{Alfen}) \exp[i(\sigma t + \phi_f)]$$

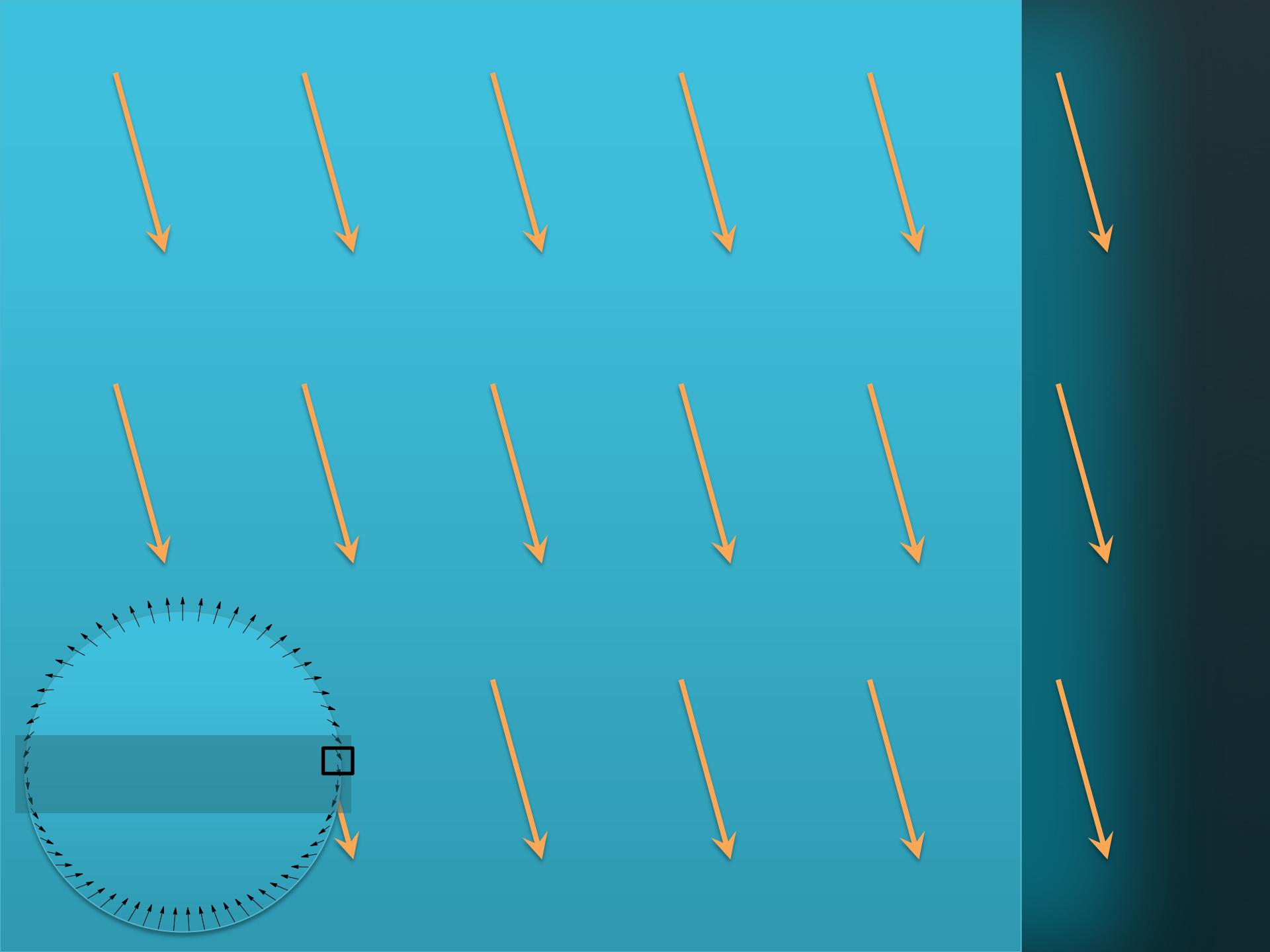
$$\xi_{\parallel} = \frac{A_s \exp[k_{\parallel} h]}{p^{1/2}} \exp[i(\sigma t + \phi_s)]$$

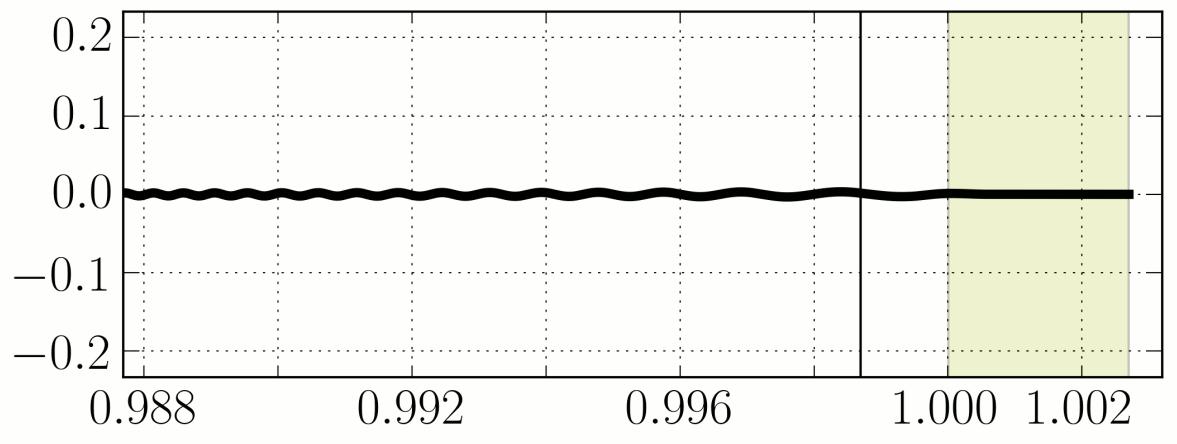
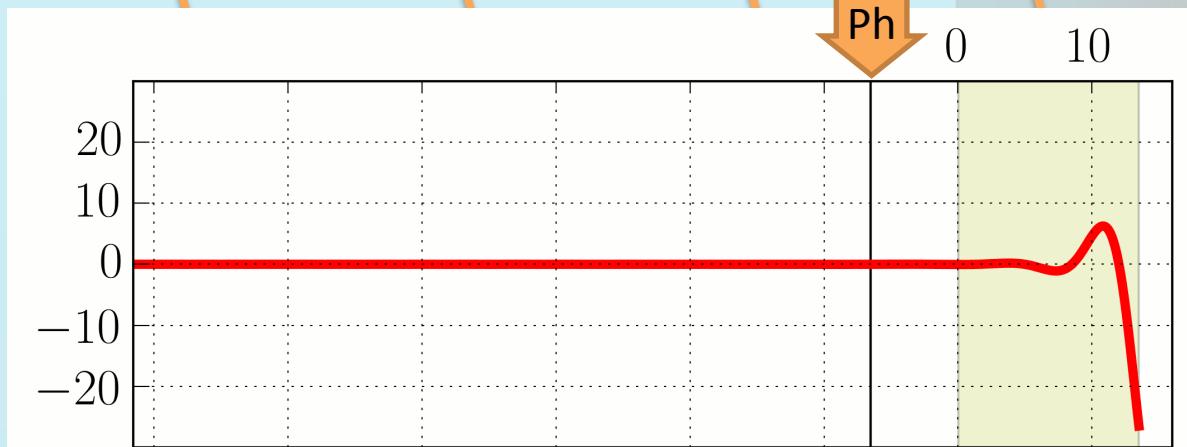
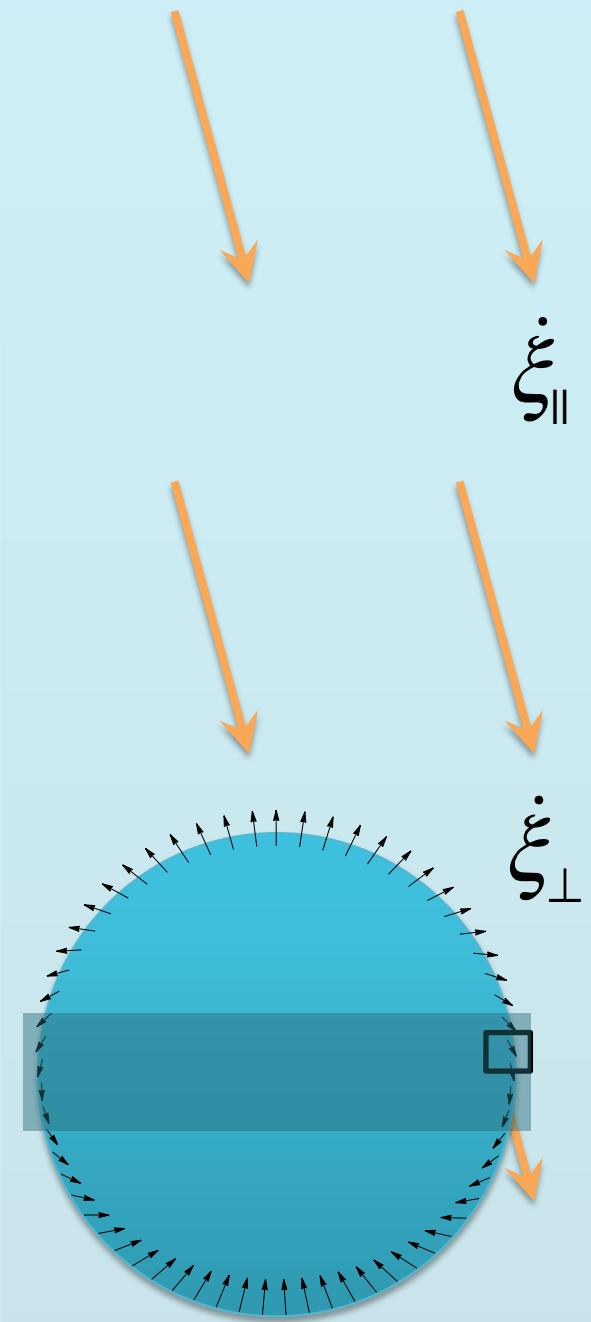
$$\xi_{\parallel} = \frac{A_s}{p^{1/2}} \exp[i(-k_{\parallel} h + \sigma t + \phi_s)]$$



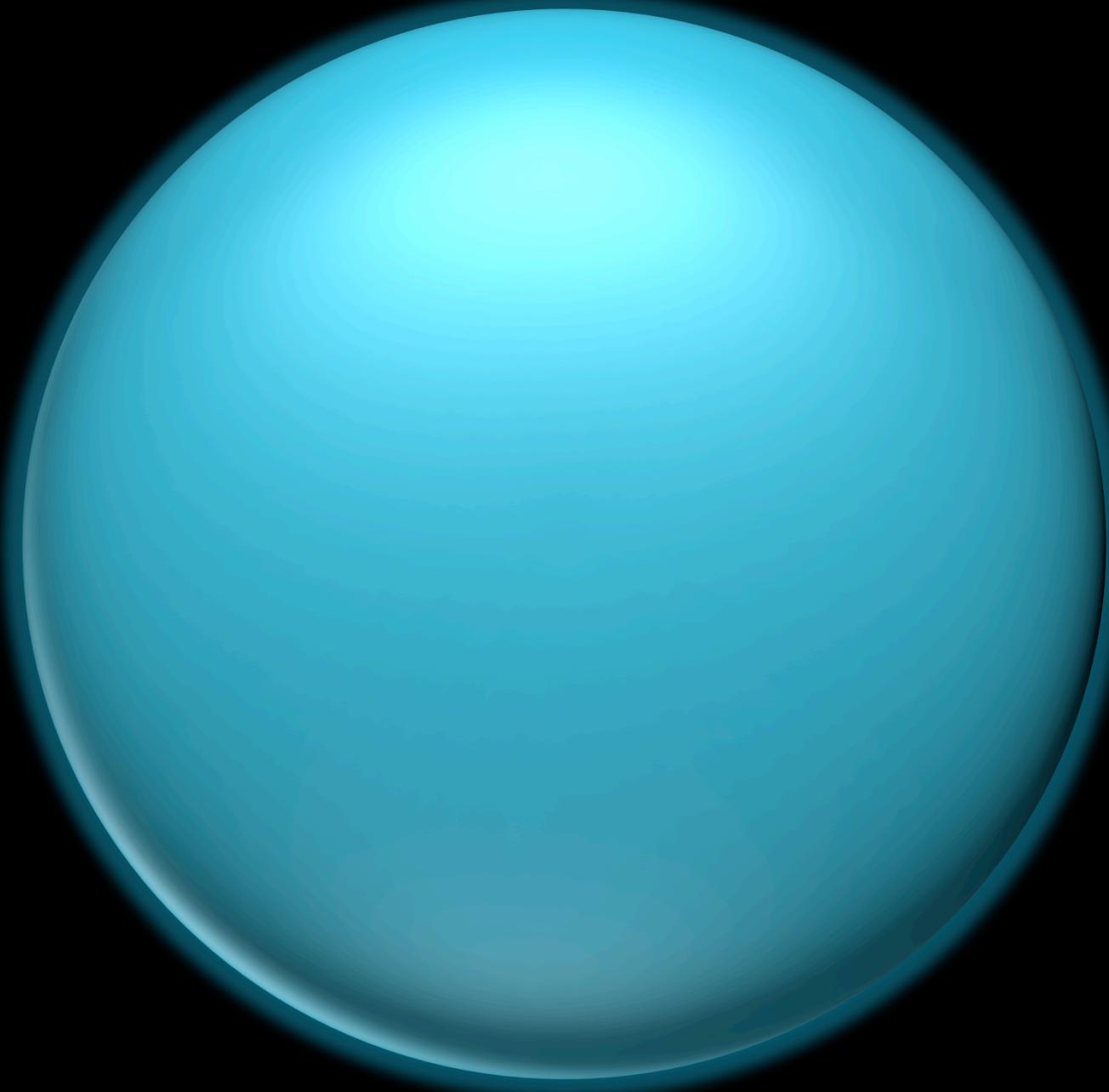






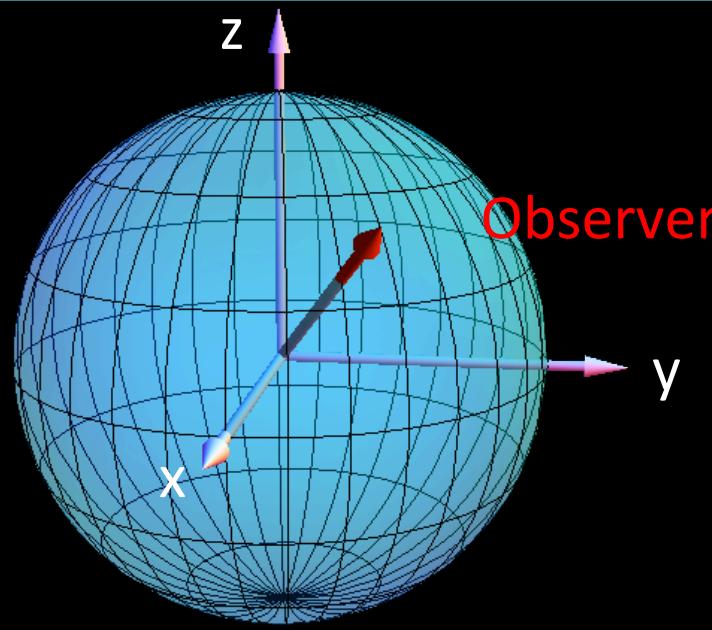


r / R



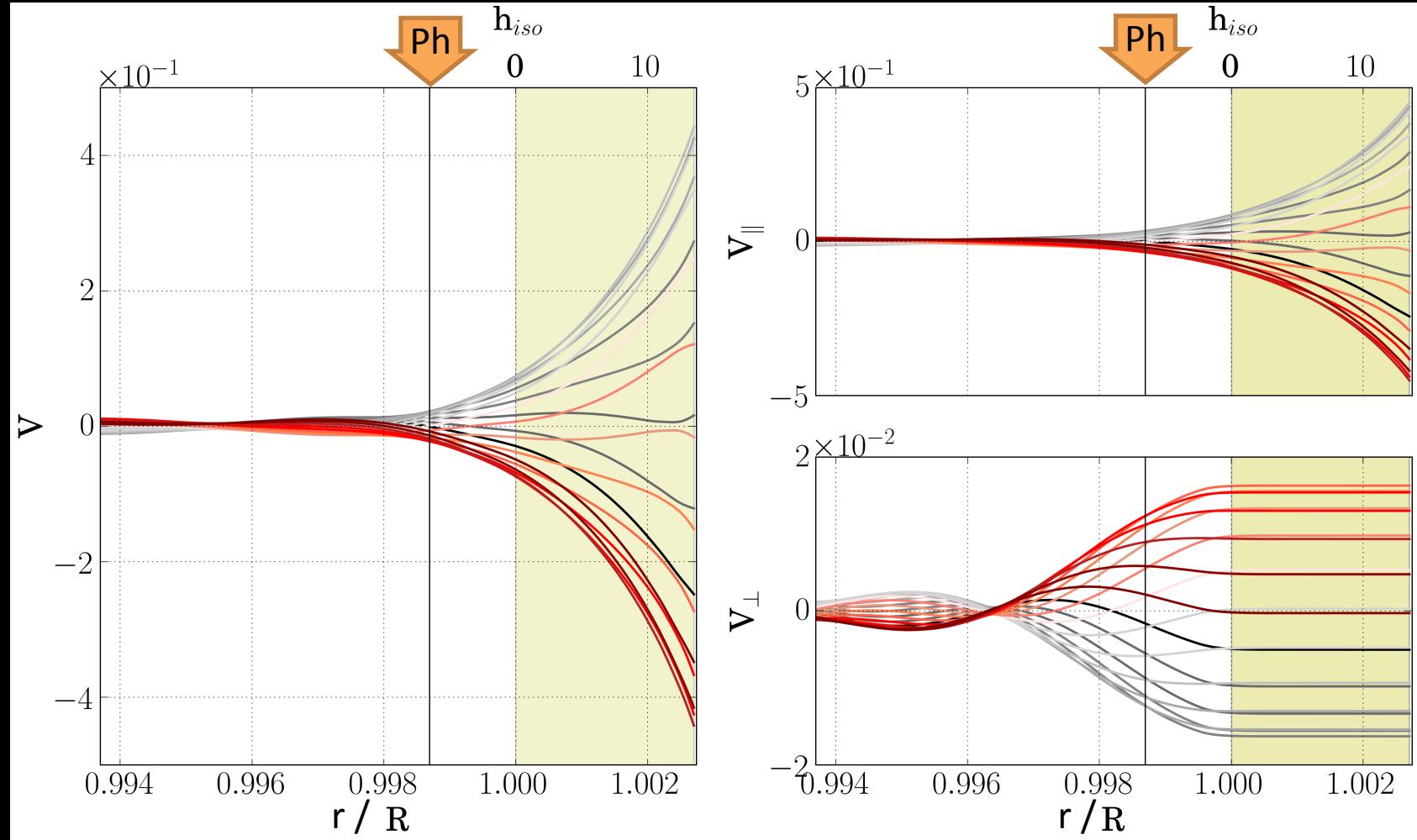
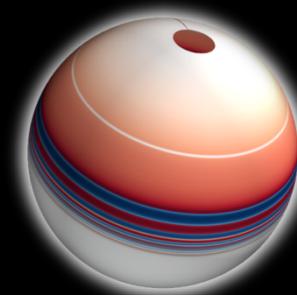
Integrated radial velocity

$$V_{\text{int}} = \int_{\varphi_i}^{\varphi_f} \int_{\theta_i}^{\theta_f} [V_z X_z + V_x X_x] \cdot \frac{1}{C_n} (1 - a(1 - \cos \theta')) \cos \theta \sin \theta d\theta d\varphi$$



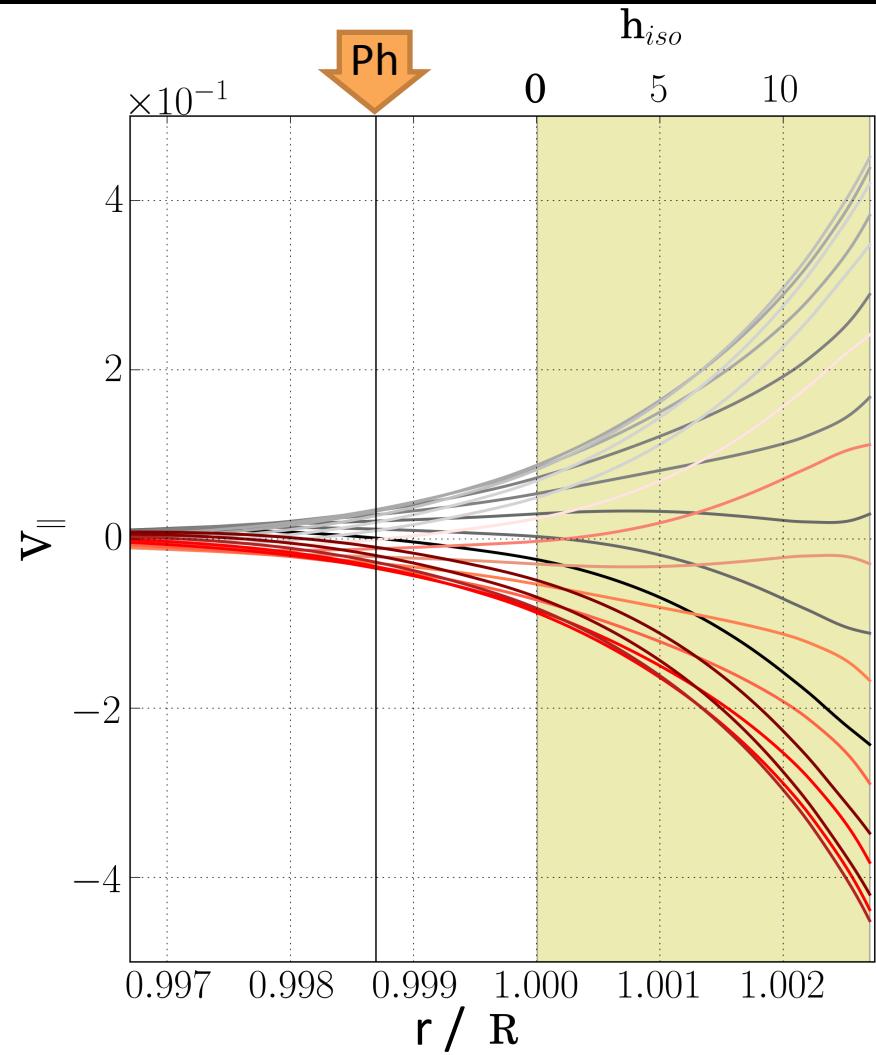
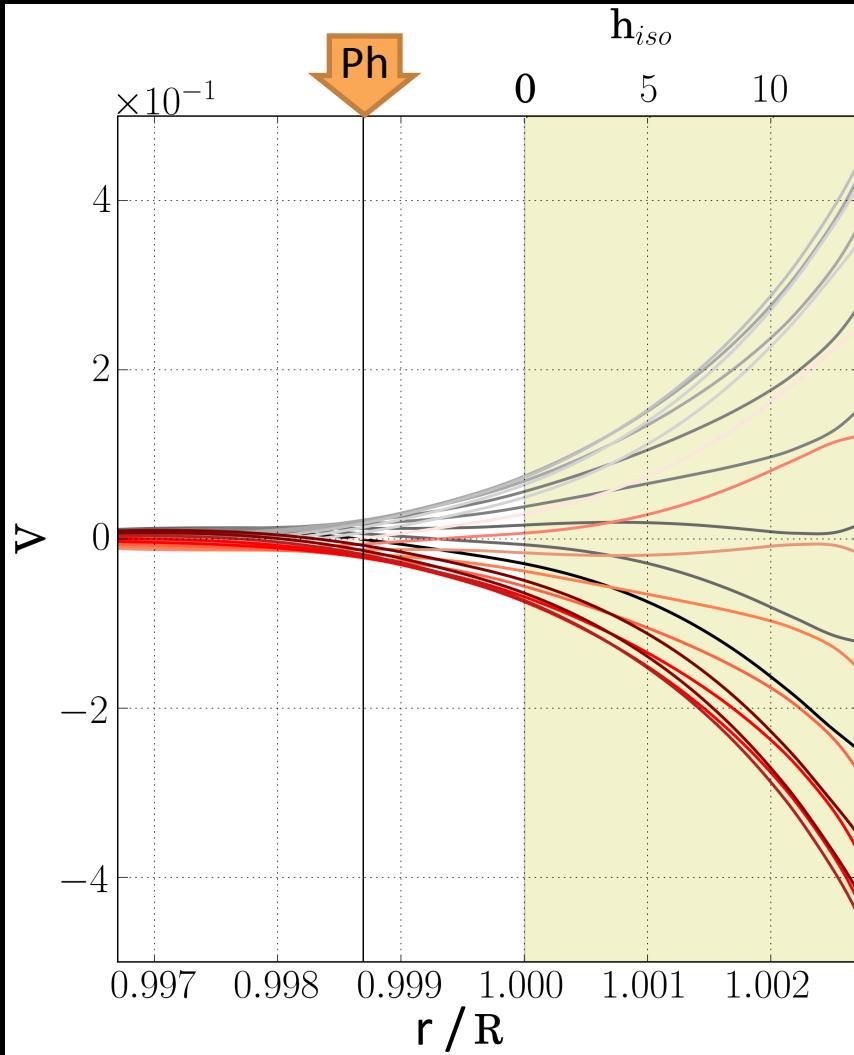
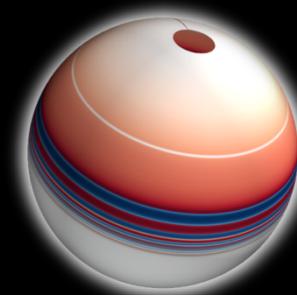
Integrated Radial Velocity

Frequency = 1.70mHz -- $\ell = 1$ -- pole-on



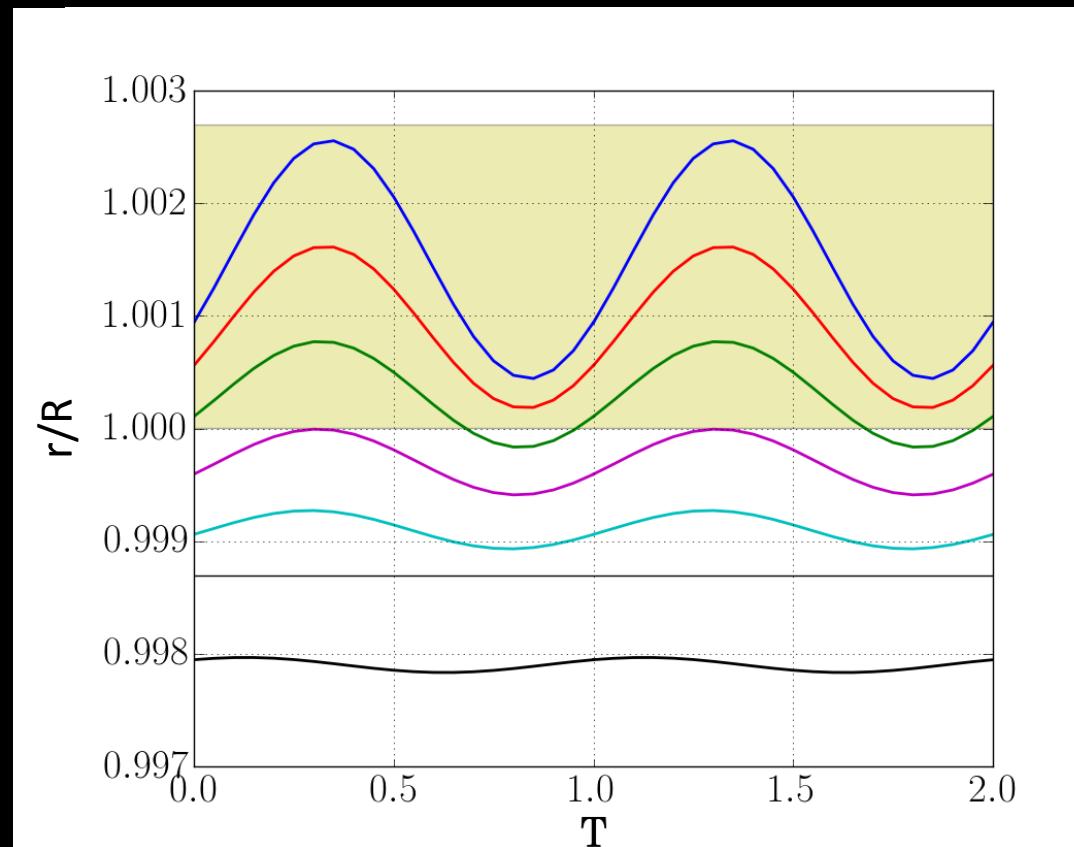
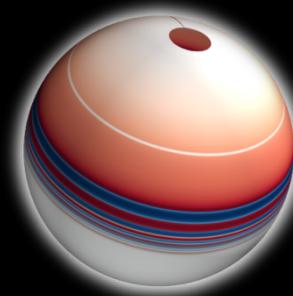
Integrated Radial Velocity

Frequency = 1.70mHz -- $\ell = 1$ -- pole-on



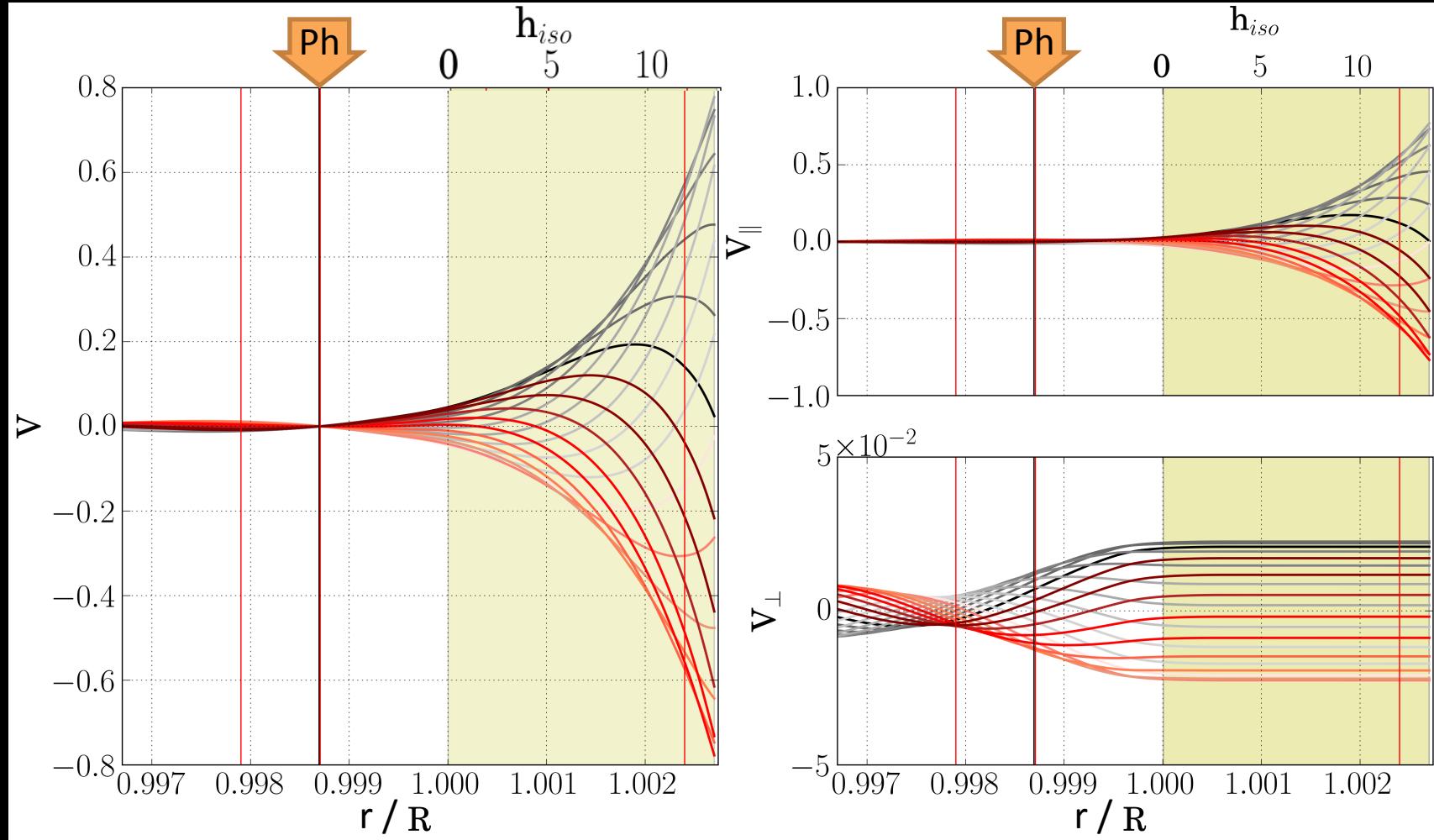
Integrated Radial Velocity

Frequency = 1.70mHz -- $\ell = 1$ -- pole-on



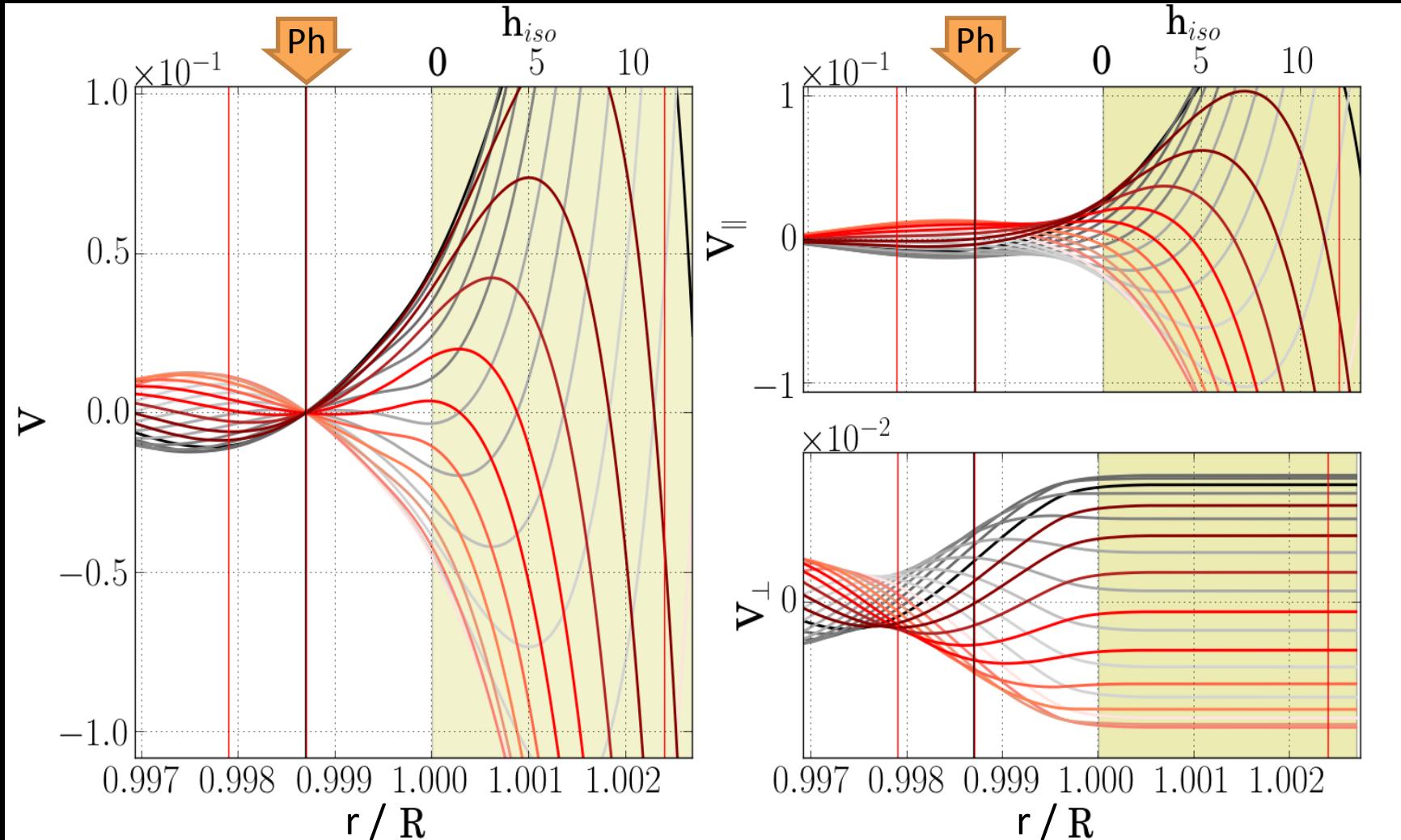
Integrated Radial Velocity

Frequency = 2.22mHz -- $I = 2$ -- equator-on



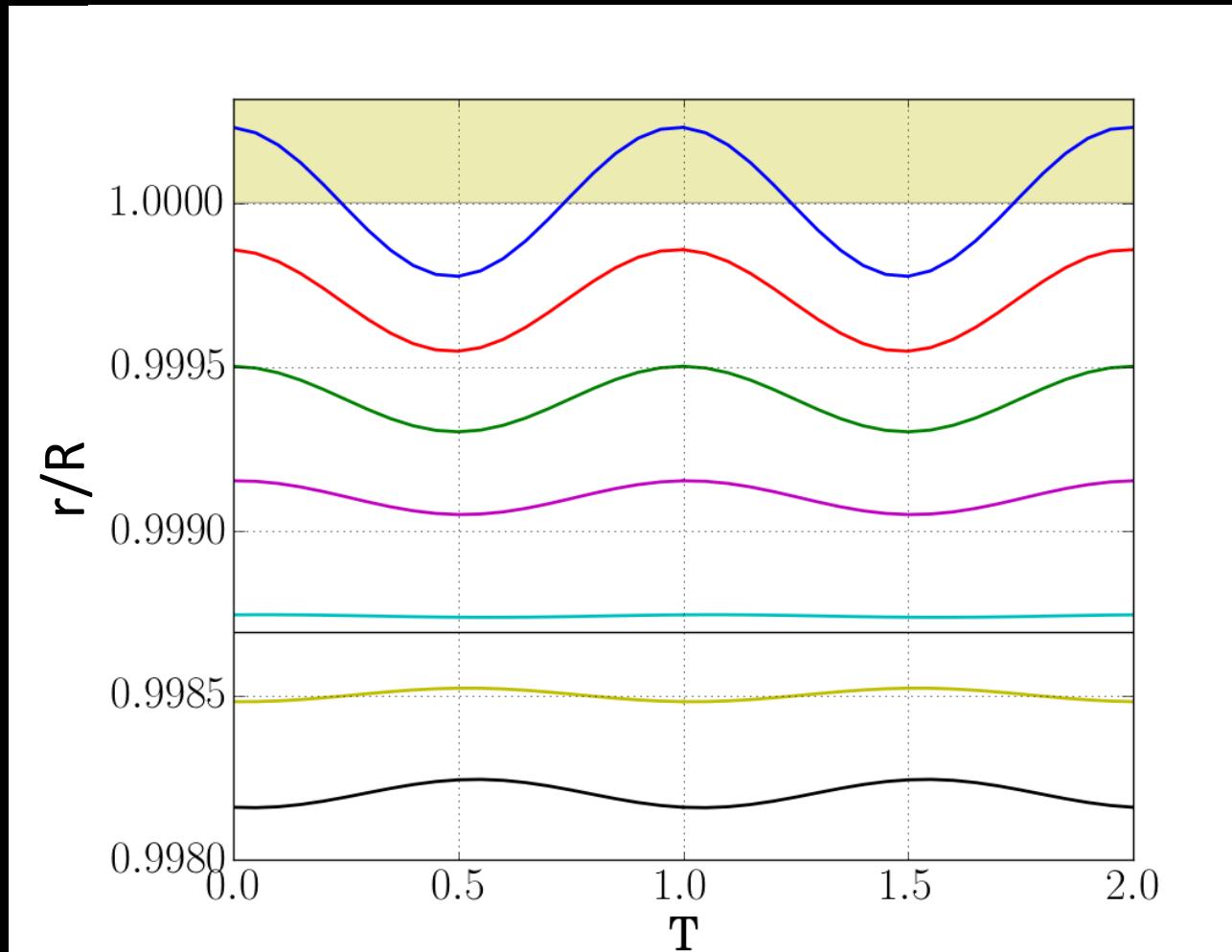
Integrated Radial Velocity

Frequency = 2.22mHz -- $l = 2$ -- equator-on



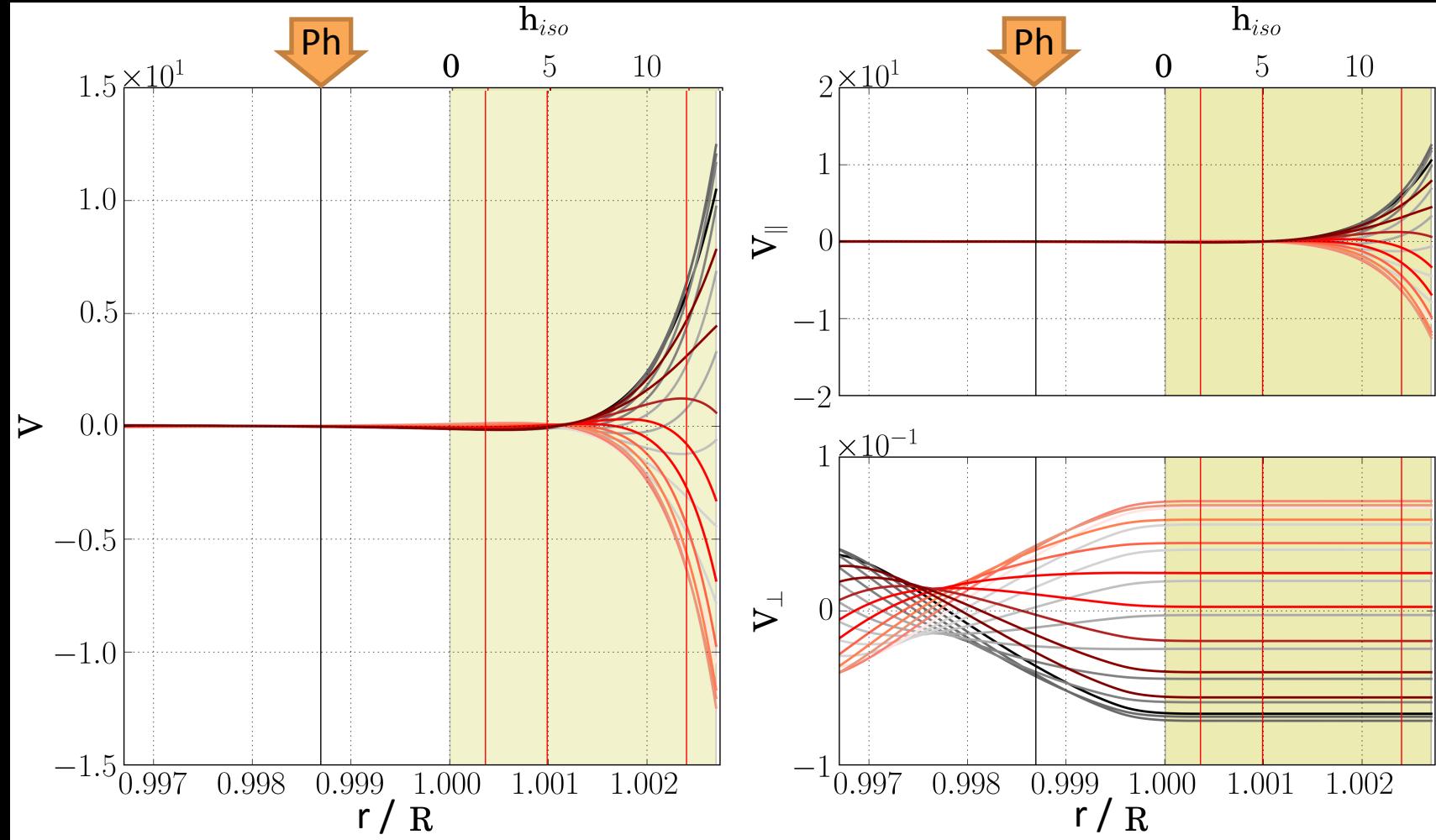
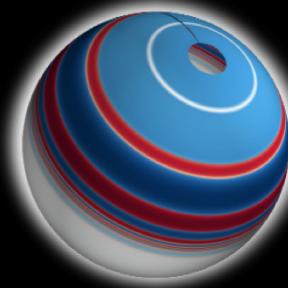
Integrated Radial Velocity

Frequency = 2.22mHz -- $I = 2$ -- equator-on



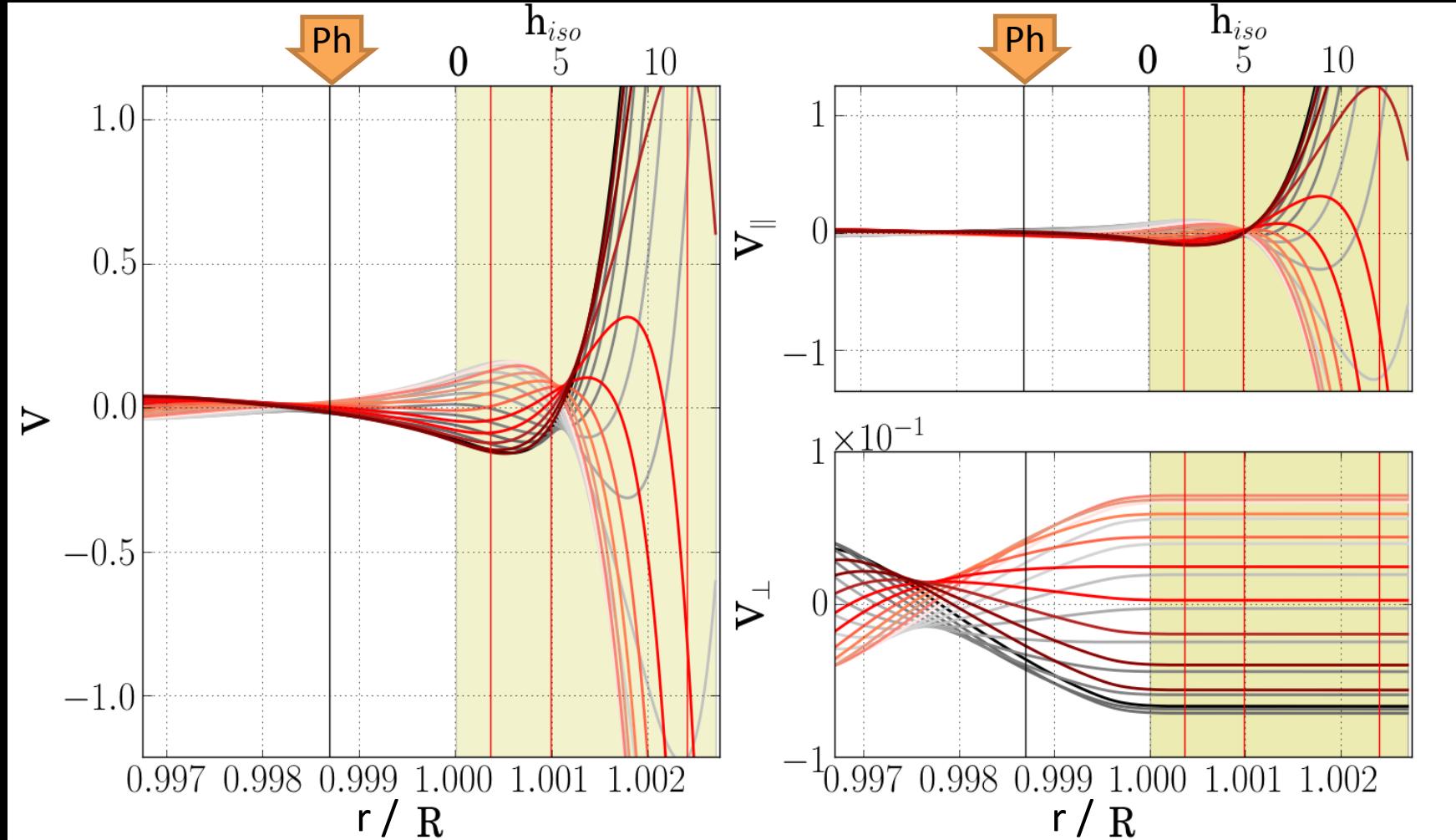
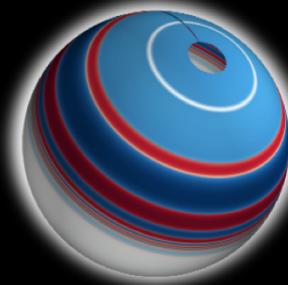
Integrated Radial Velocity

Frequency = 2.4mHz -- $I = 4$ -- pole-on



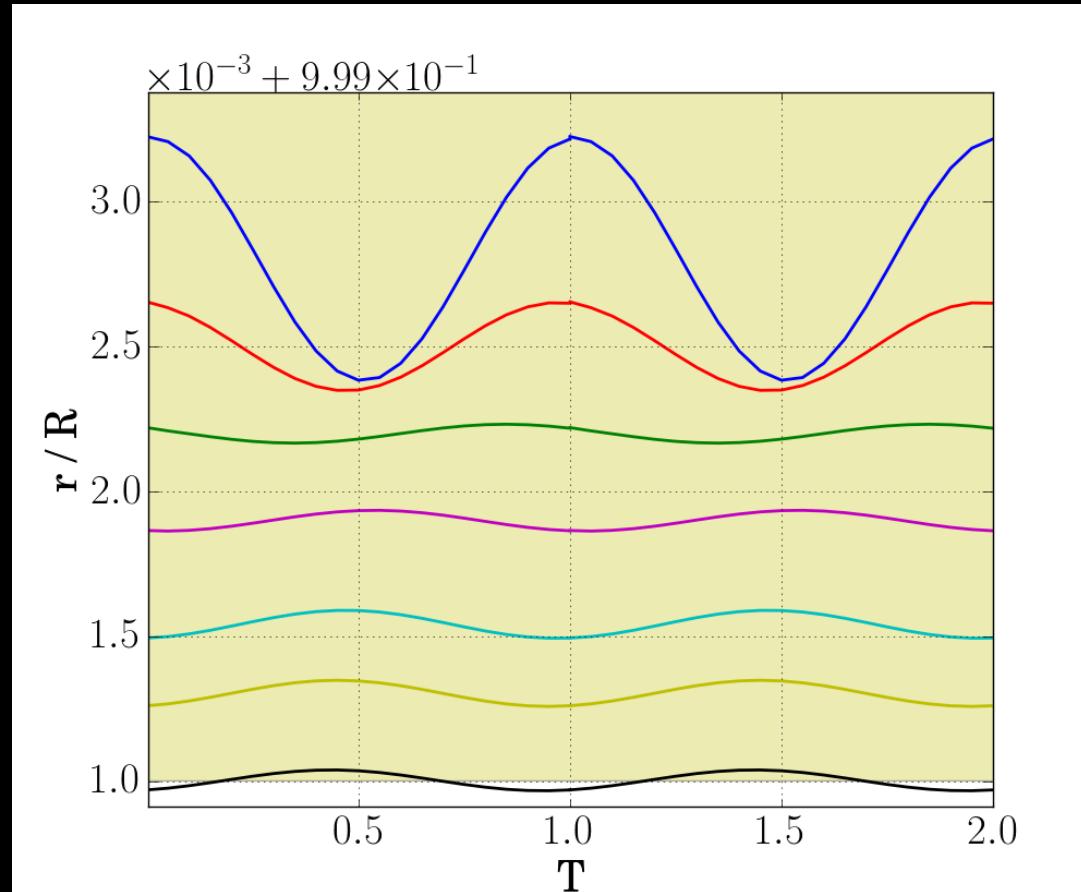
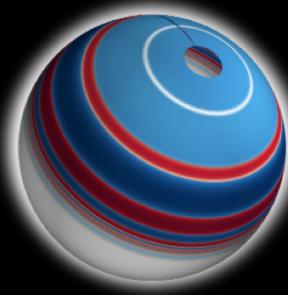
Integrated Radial Velocity

Frequency = 2.4mHz -- $|l| = 4$ -- pole-on



Integrated Radial Velocity

Frequency = 2.4mHz -- $I = 4$ -- pole-on



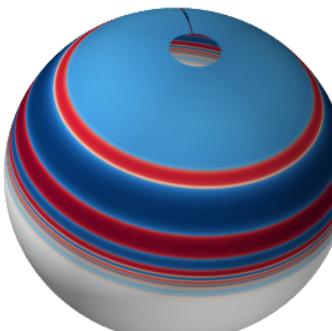
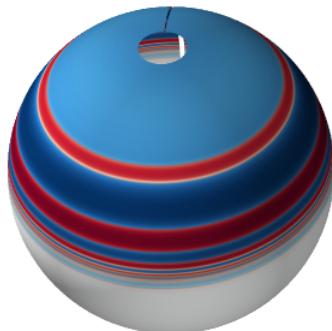
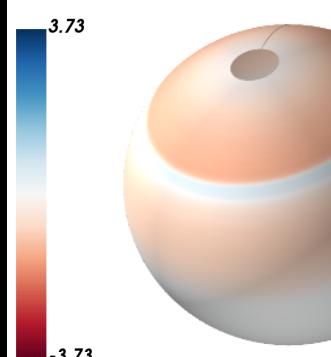
Summary:

- We built a tool to calculate the Integrated radial velocities
- We have confirmed that the observer sees a very complex picture due to the fact that there is an integration over the disc
- We plan to use the tool to forward model the radial velocity observed in roAp stars

Thanks

Paola.Quitral@astro.up.pt

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