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The UCD/cE Divide: Aiming to Bridge the Gap

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MPIA

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THE UNIVERSITY
of NORTH CAROLINA
at CHAPEL HILL



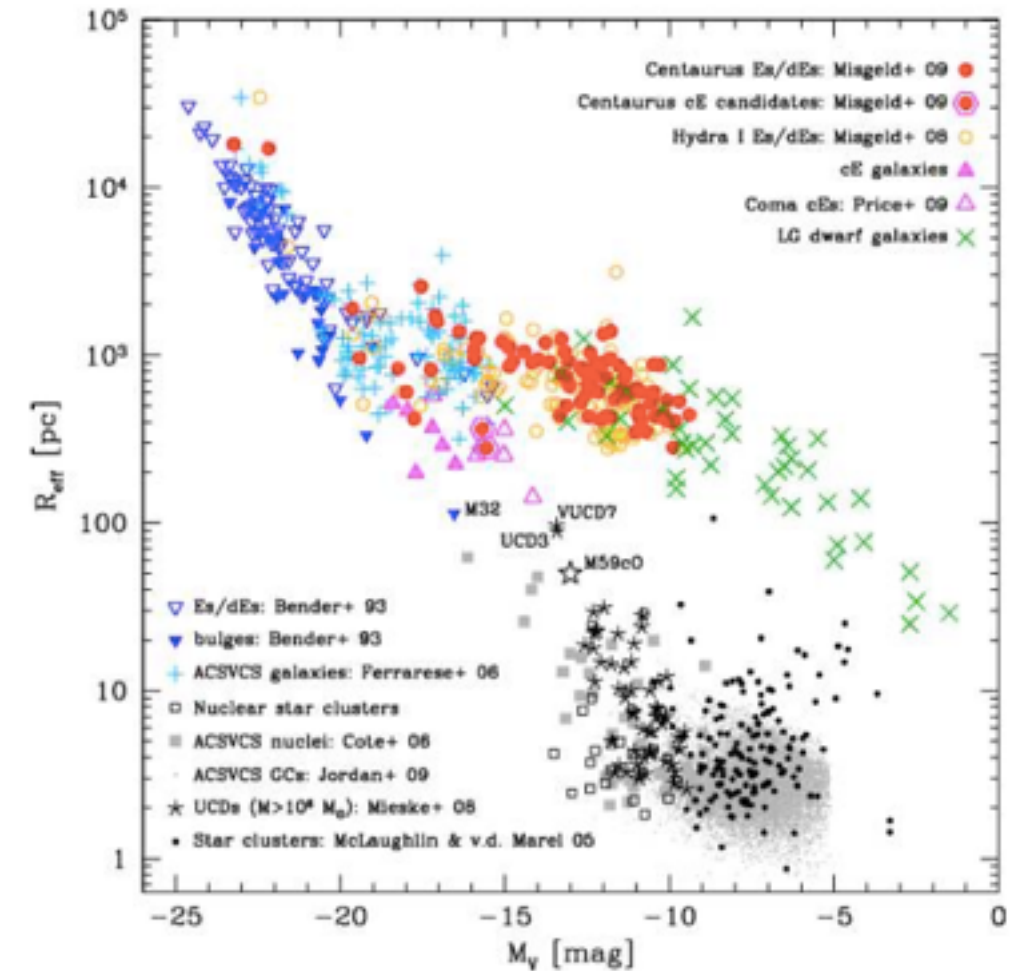
Motivation

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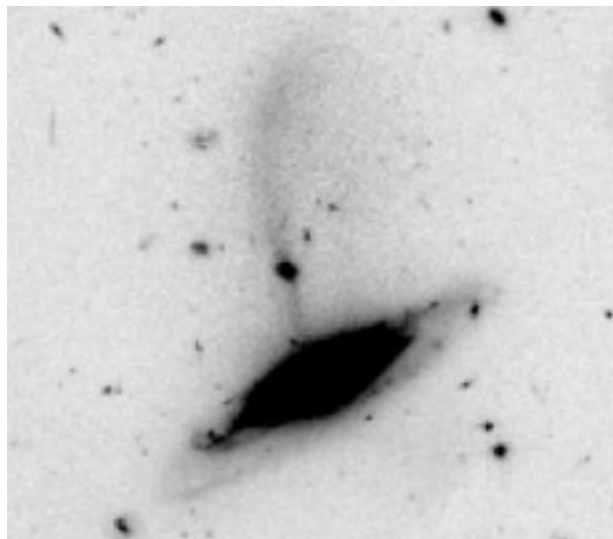
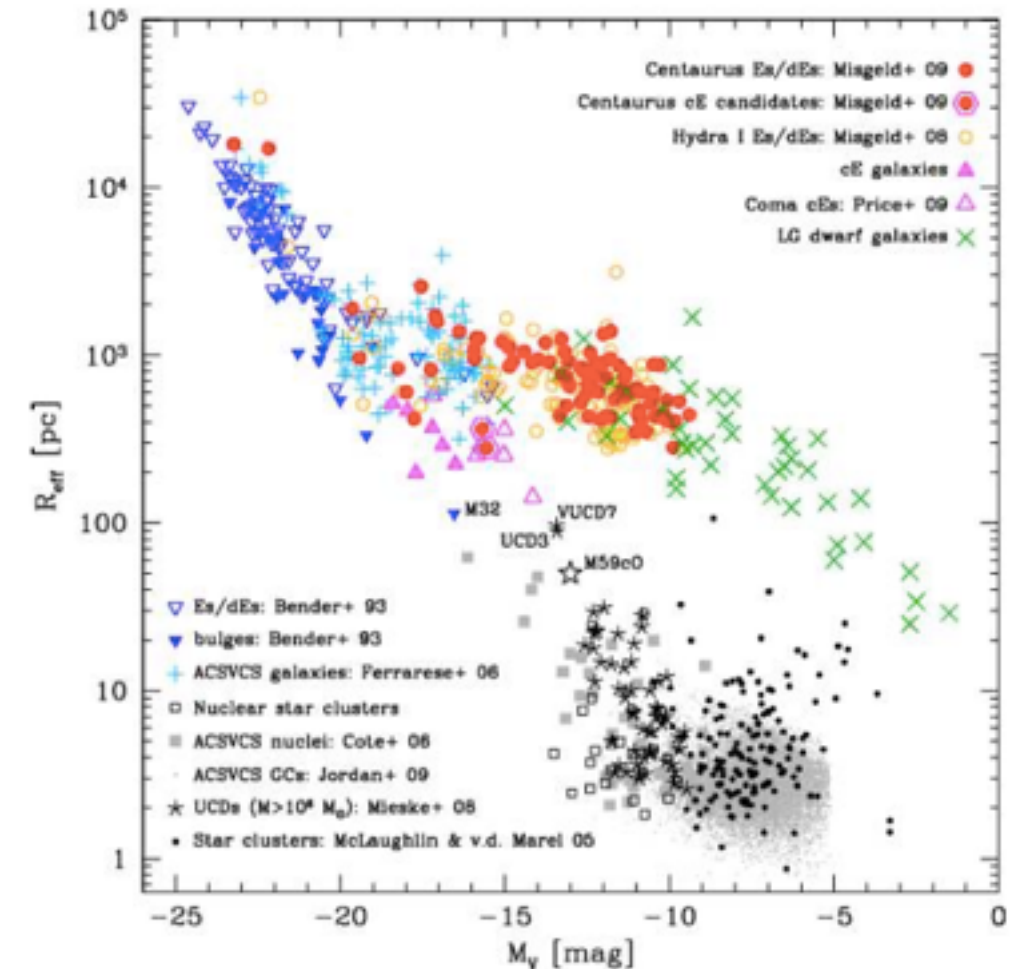
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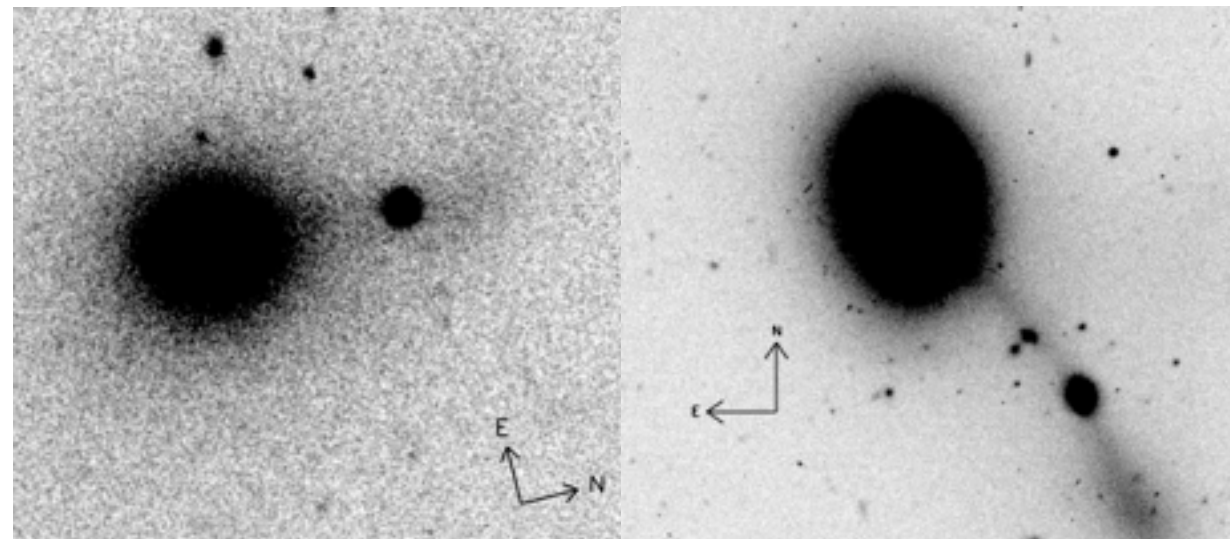
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- 2) They are thought to form through the stripping of larger galaxies - potentially long lived signs of merger histories.



Forbes et al. 2003

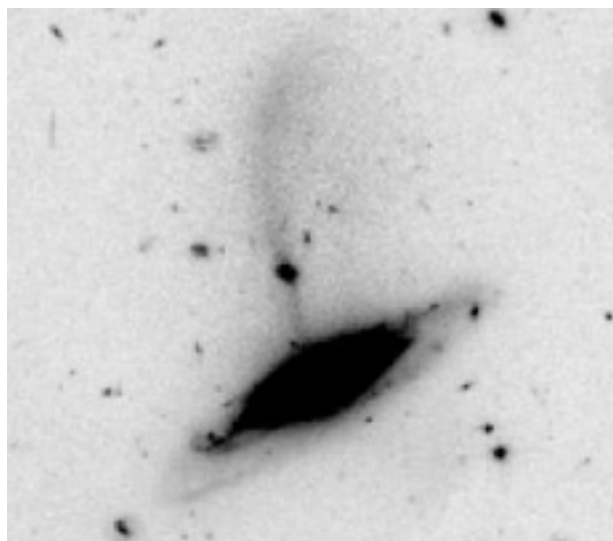
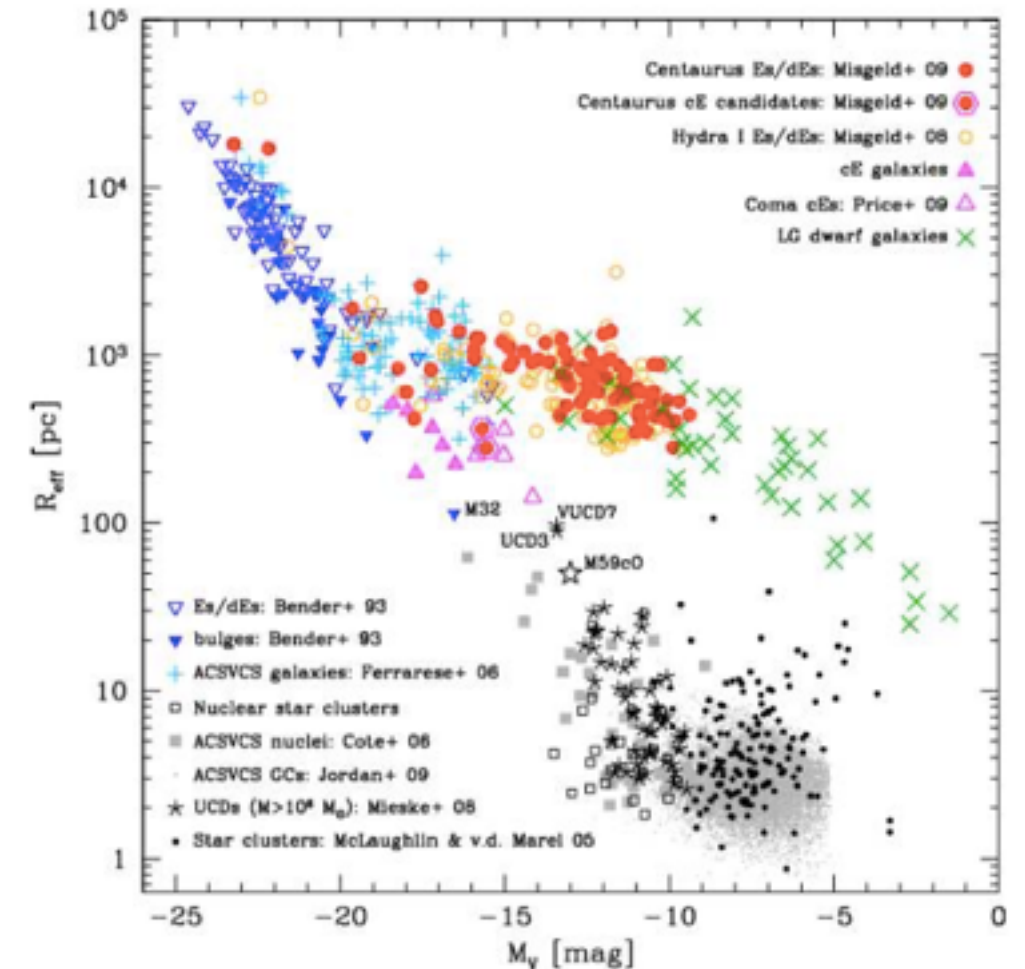


Huxor, Phillips, Price & Harniman 2011

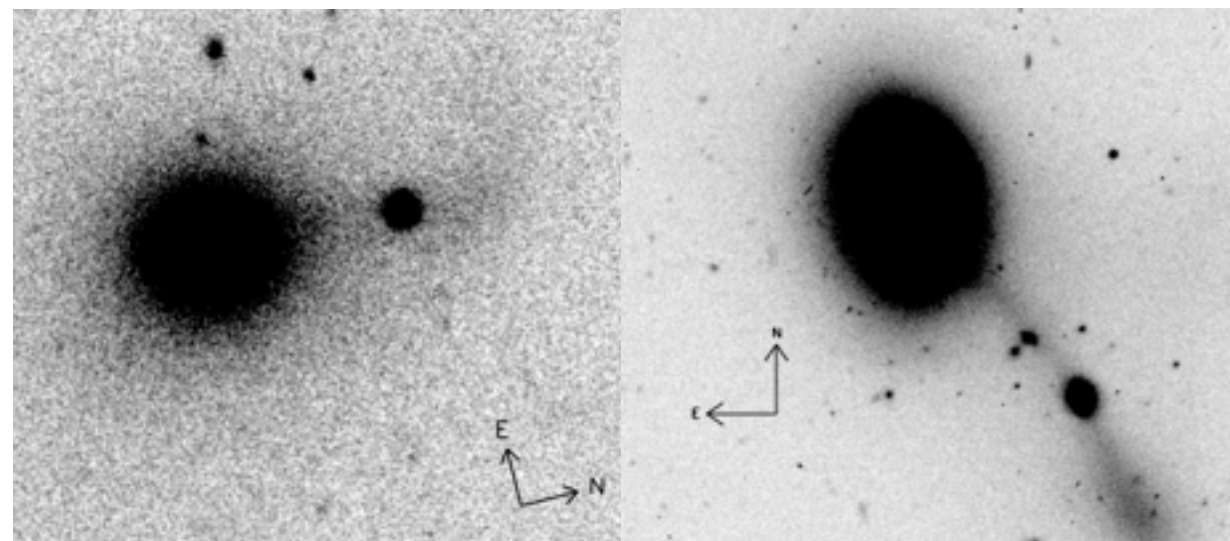
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cEs are interesting.

- 1) They are intermediate between star clusters and galaxies in most properties.
- 2) They are thought to form through the stripping of larger galaxies - potentially long lived signs of merger histories. - [though see talk by Analia Smith-Castelli and Huxor et al. 2013 for other possibilities.](#)



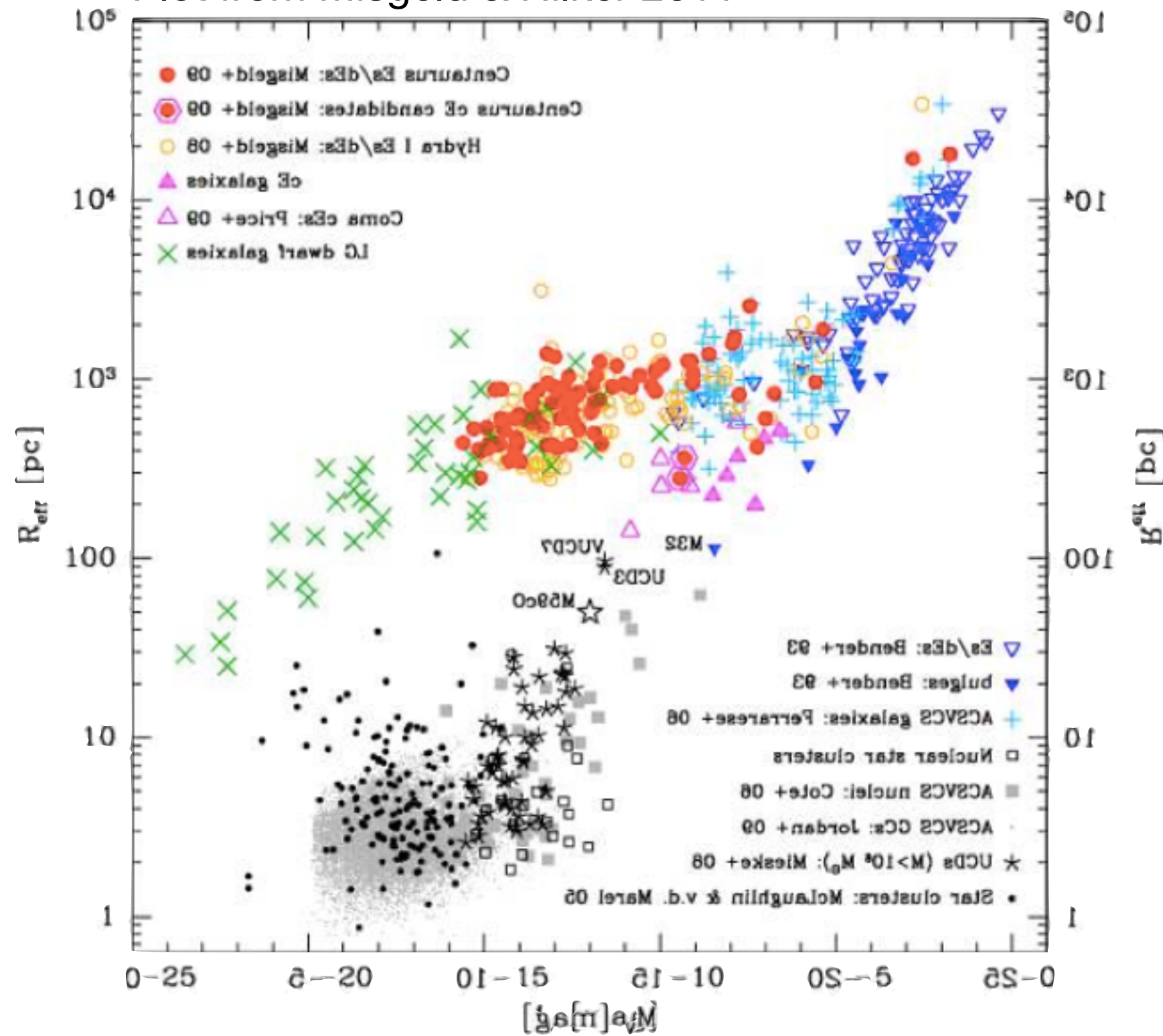
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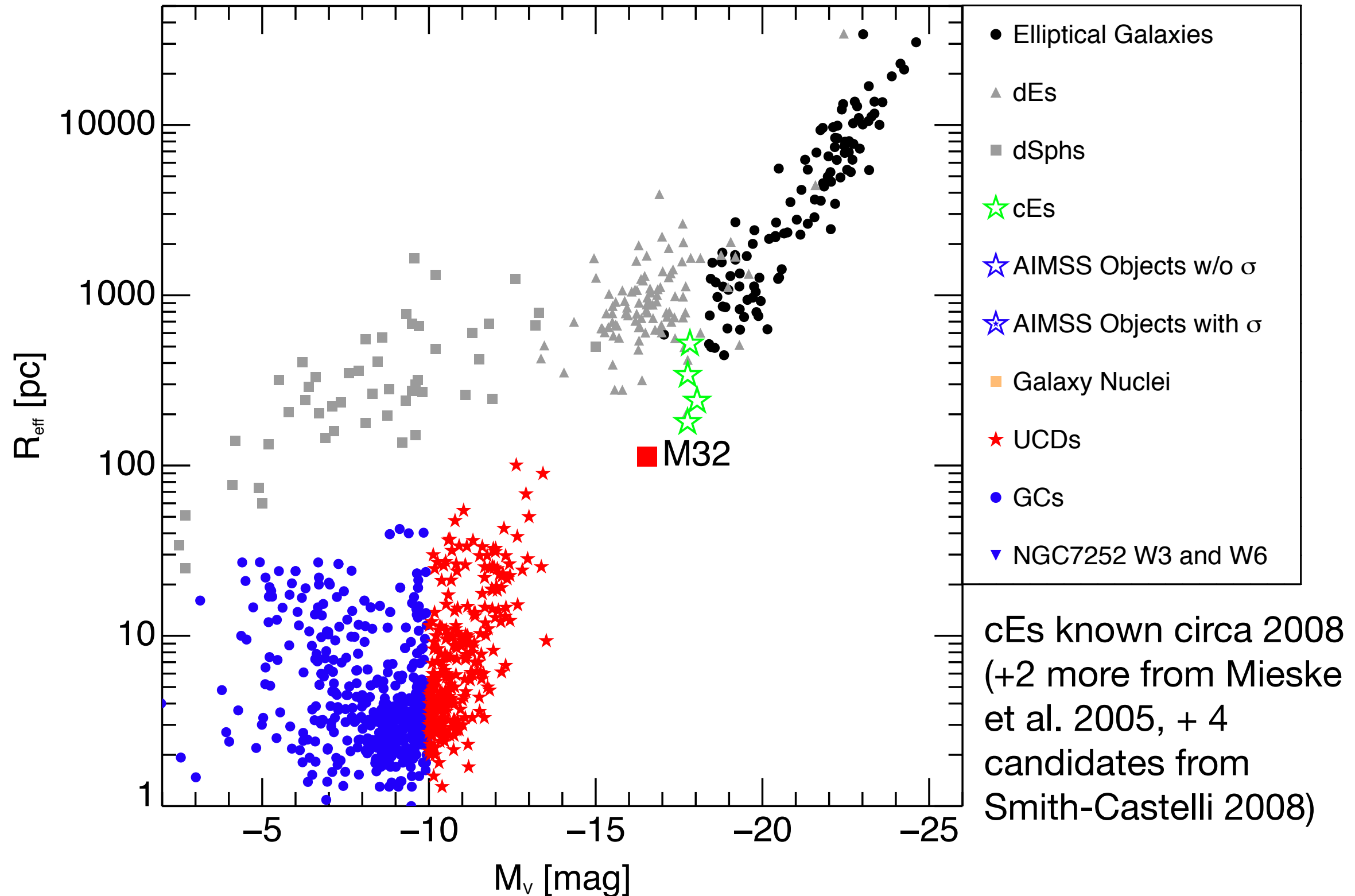
cE Recent History

Plot from Misgeld & Hilker 2011



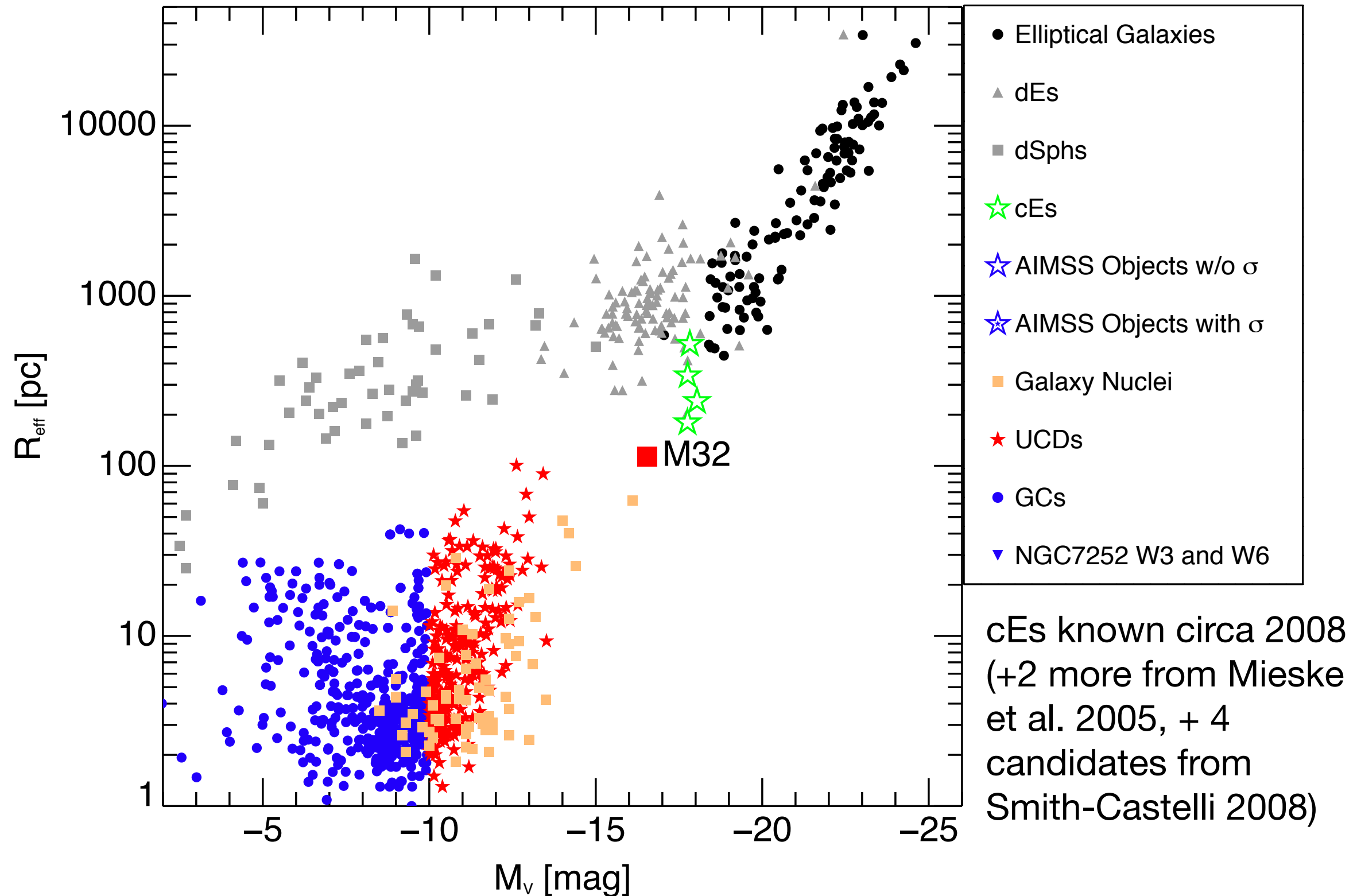
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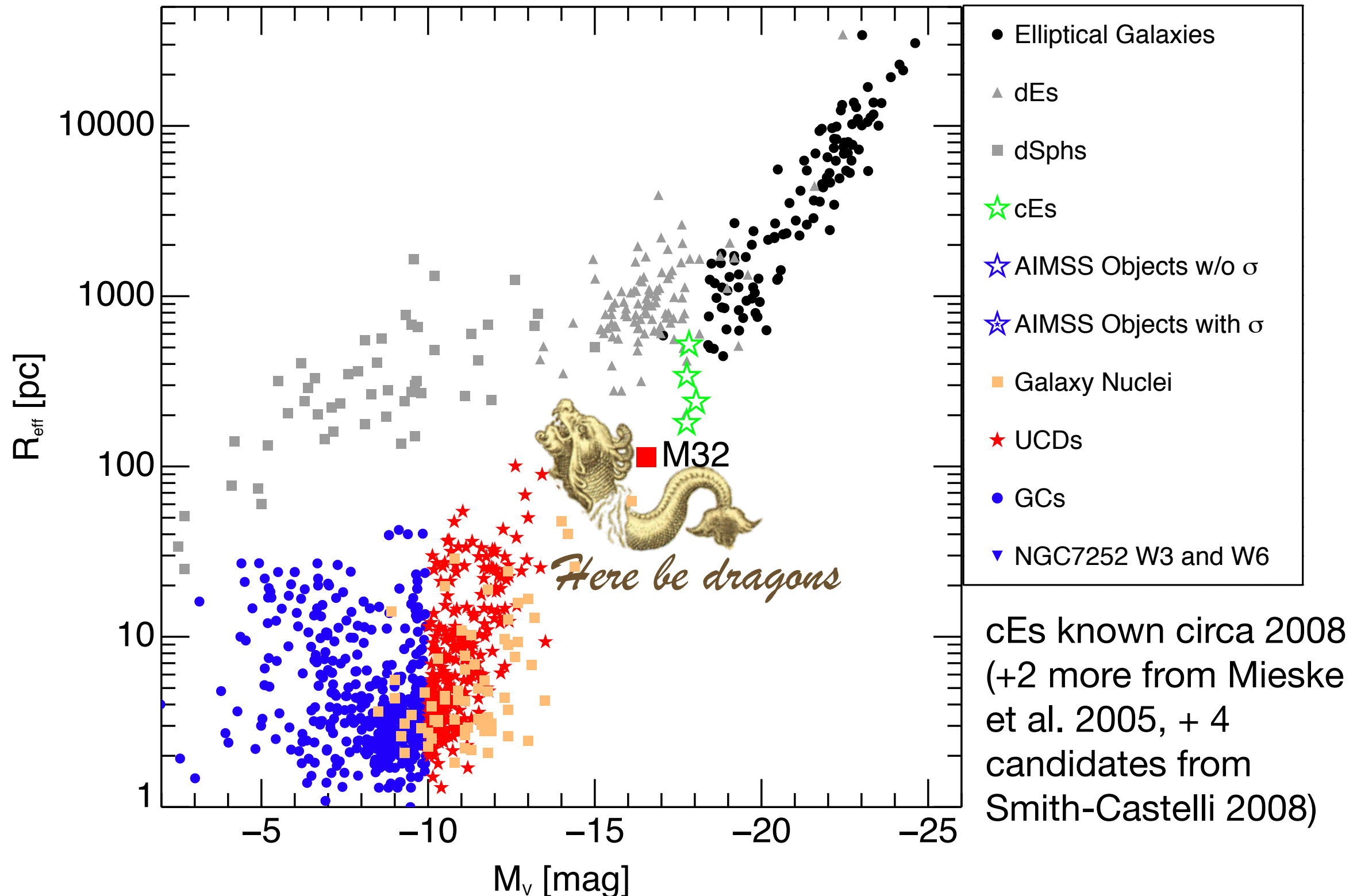
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The AIMSS Survey

The Archive of Intermediate Mass Stellar Systems - a two part project:

1.) The AIMSS Survey - Archival HST survey for compact stellar systems: massive GCs, UCDs, and cEs.

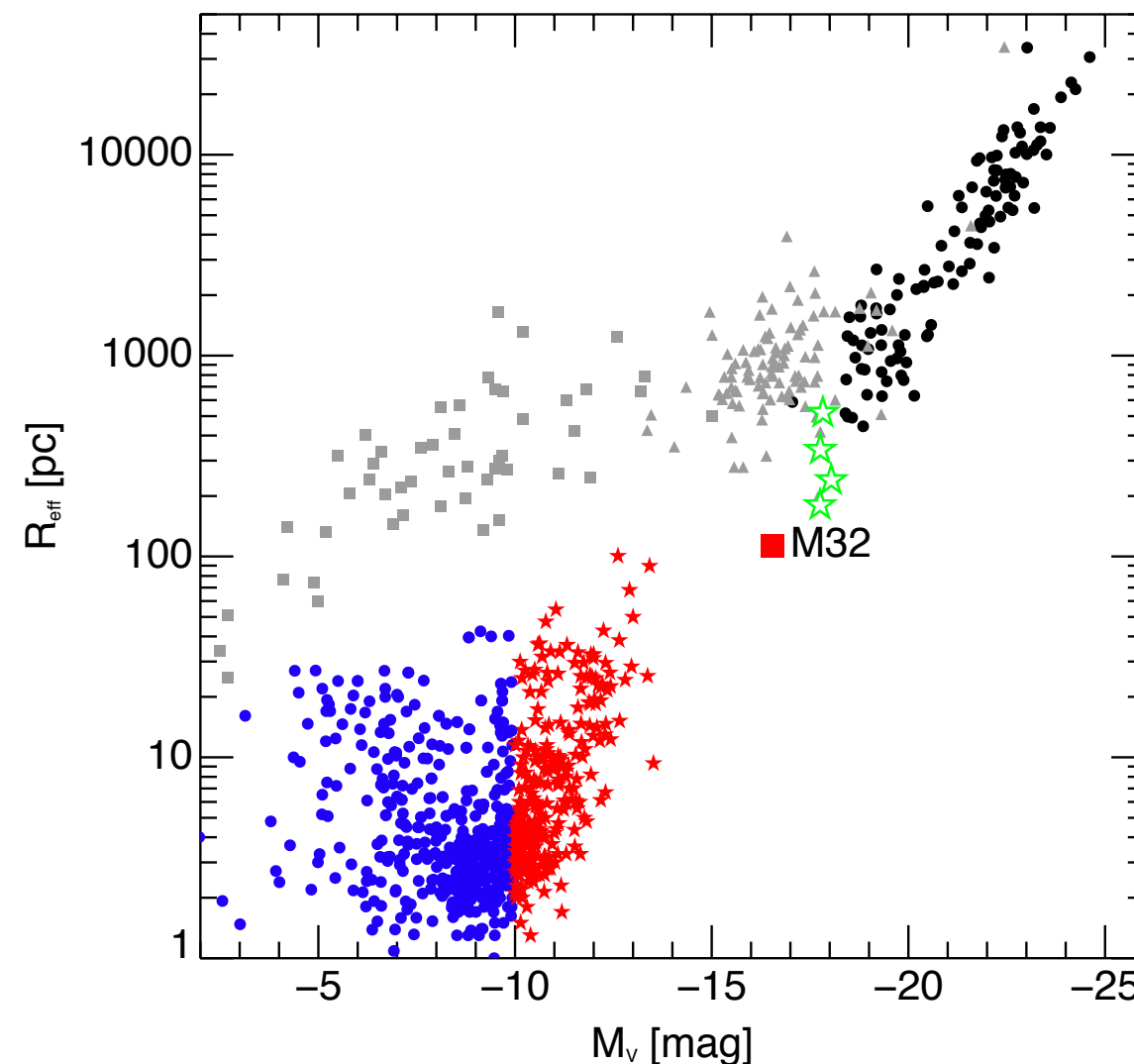
Select candidates based on position in luminosity-size relation assuming they were at the distance of the host galaxy:

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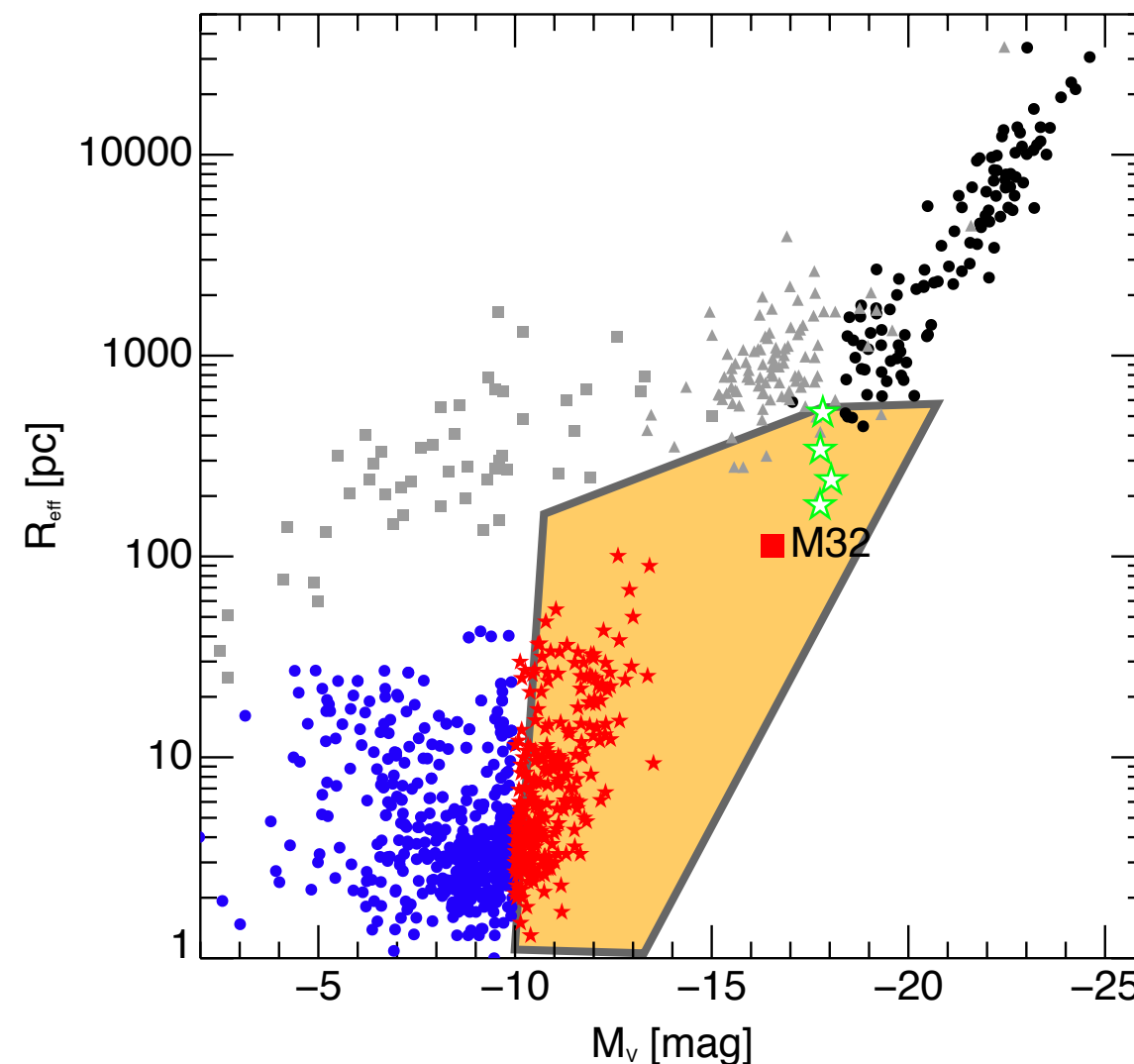


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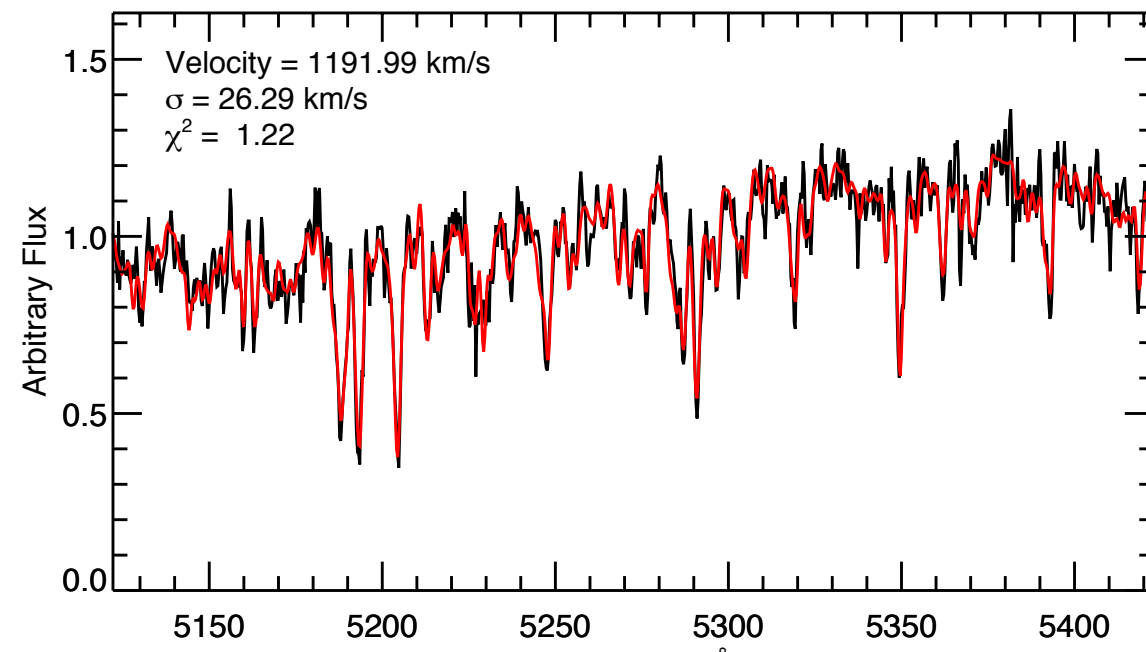
The AIMSS Survey



Confirming redshifts from Keck, SOAR, and Gemini + internal velocity dispersions for most



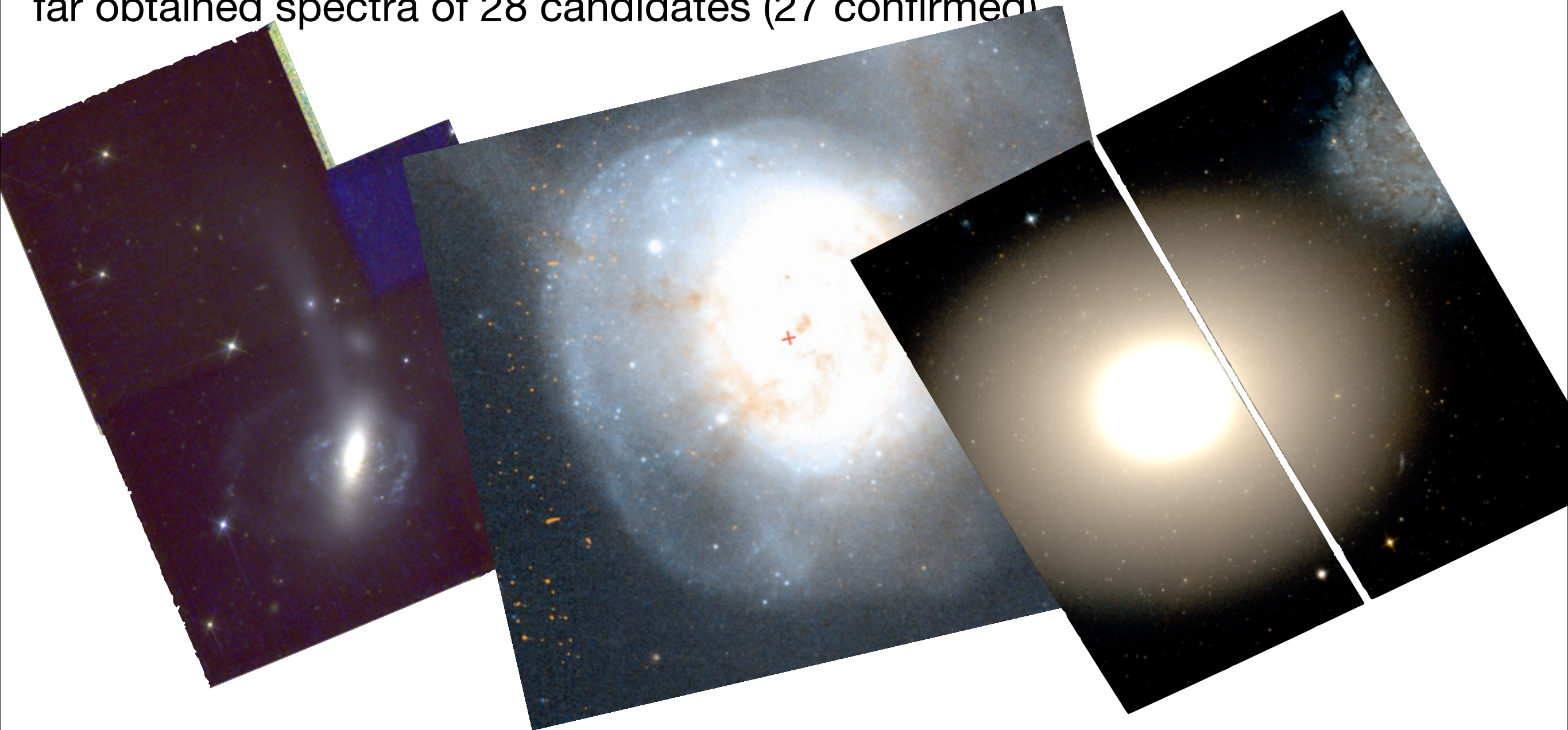
NGC4546_UCD1



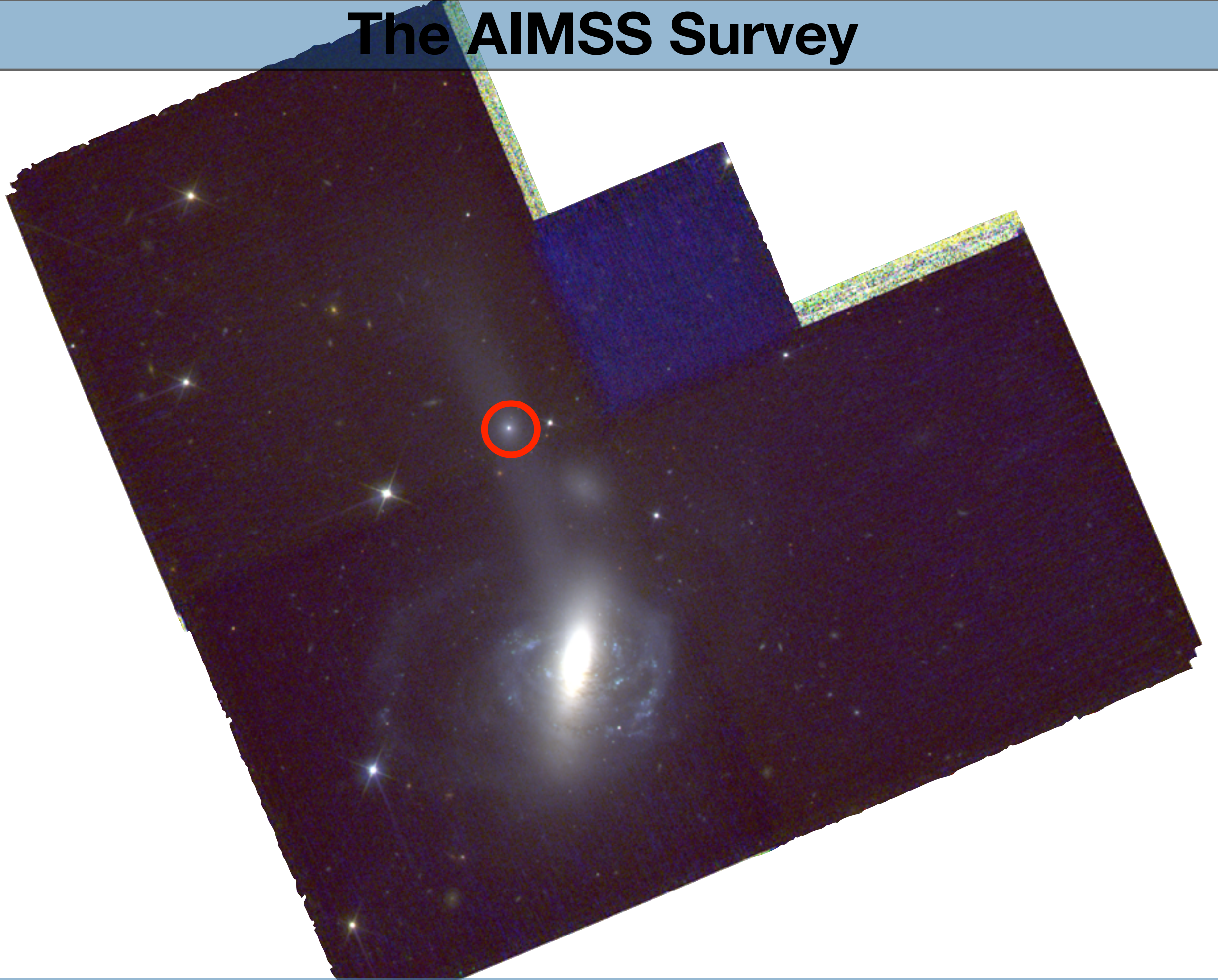
The AIMSS Survey

Initially searched HST imaging of 76 nearby galaxies - found 9 UCD candidates - obtained spectra of 4 (all 4 confirmed). Published in Norris & Kannappan 2011.

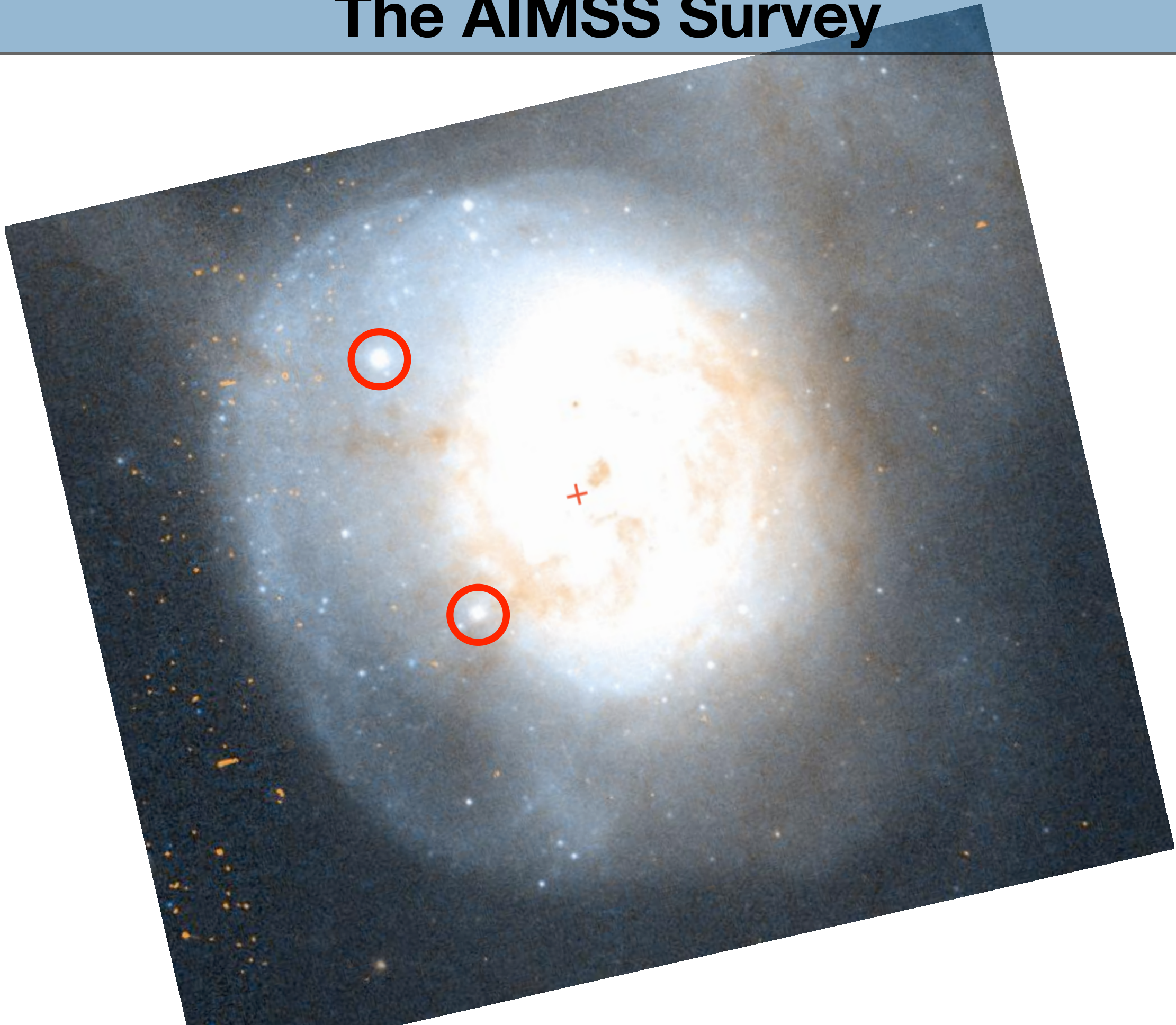
Then expanded to HST imaging of >500 galaxies - hundreds of candidates, so far obtained spectra of 28 candidates (27 confirmed)



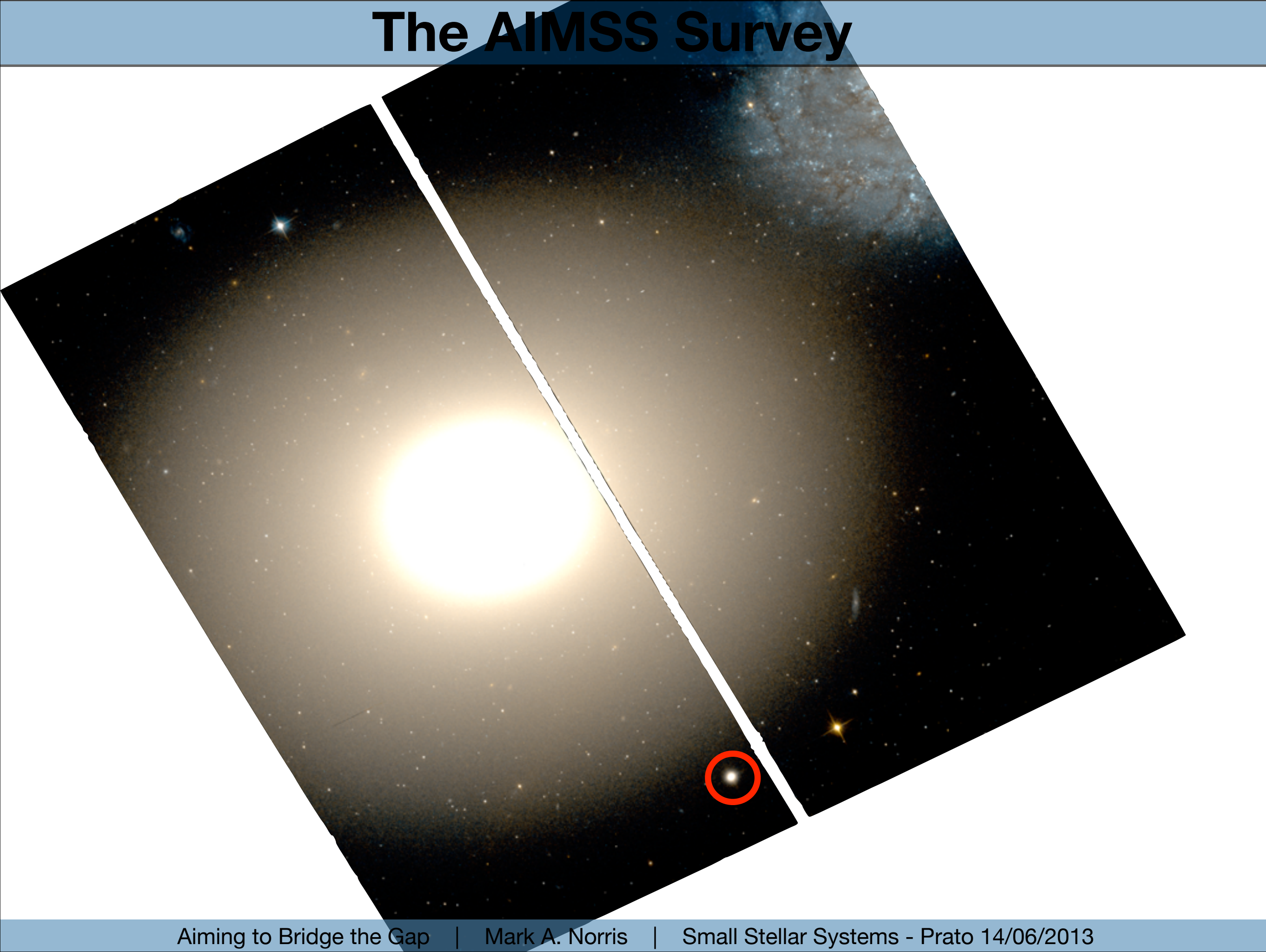
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Strengths: Sensitive to all objects - GCs, YMCs, UCDs, cEs
Structural information, colour gradients

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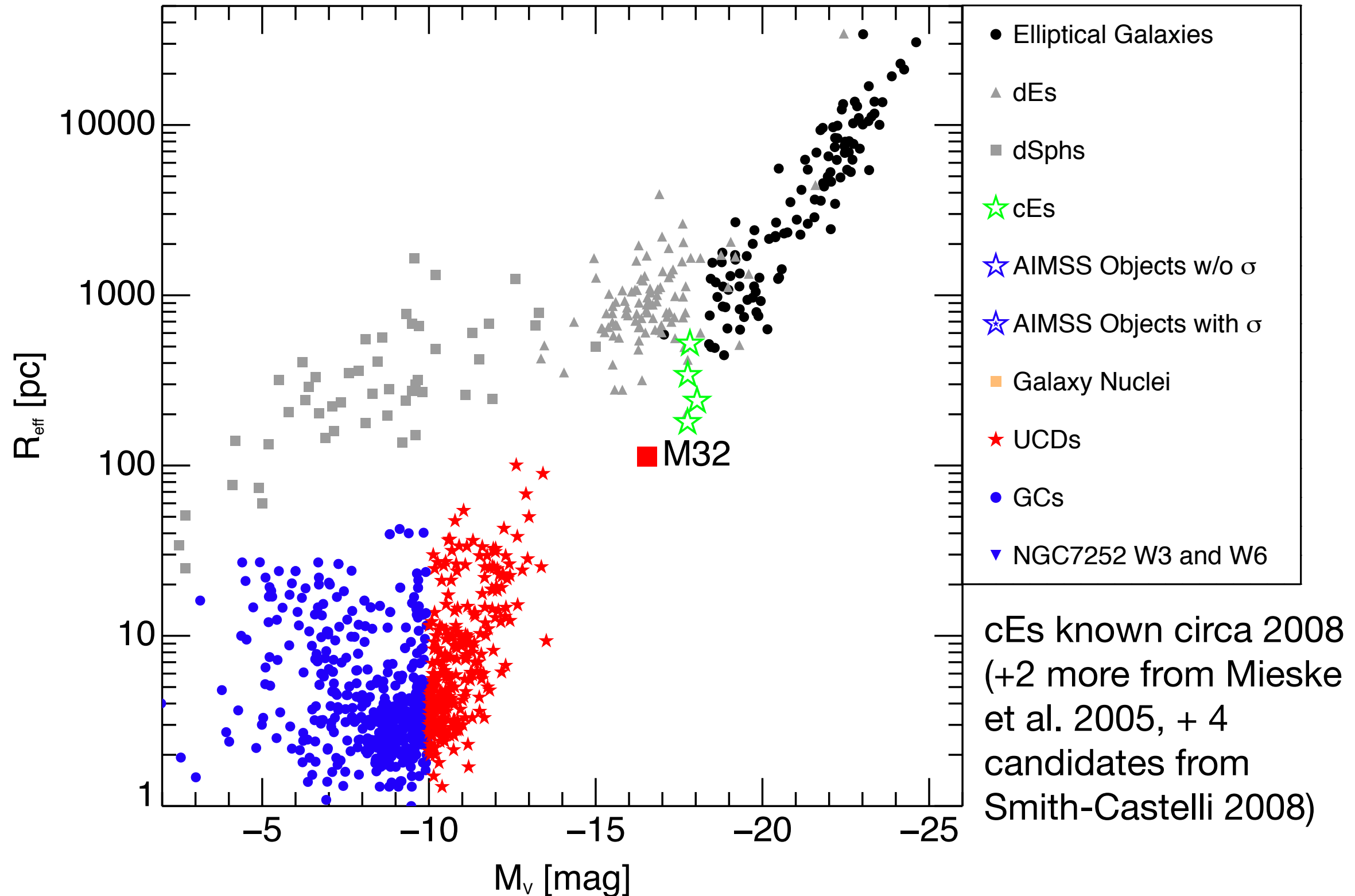
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Weaknesses: Totally inhomogeneous - get whatever HST observed
Environmental analyses ~impossible - need volume/magnitude
limited spectroscopic samples [see talk by Igor Chilingarian next](#)

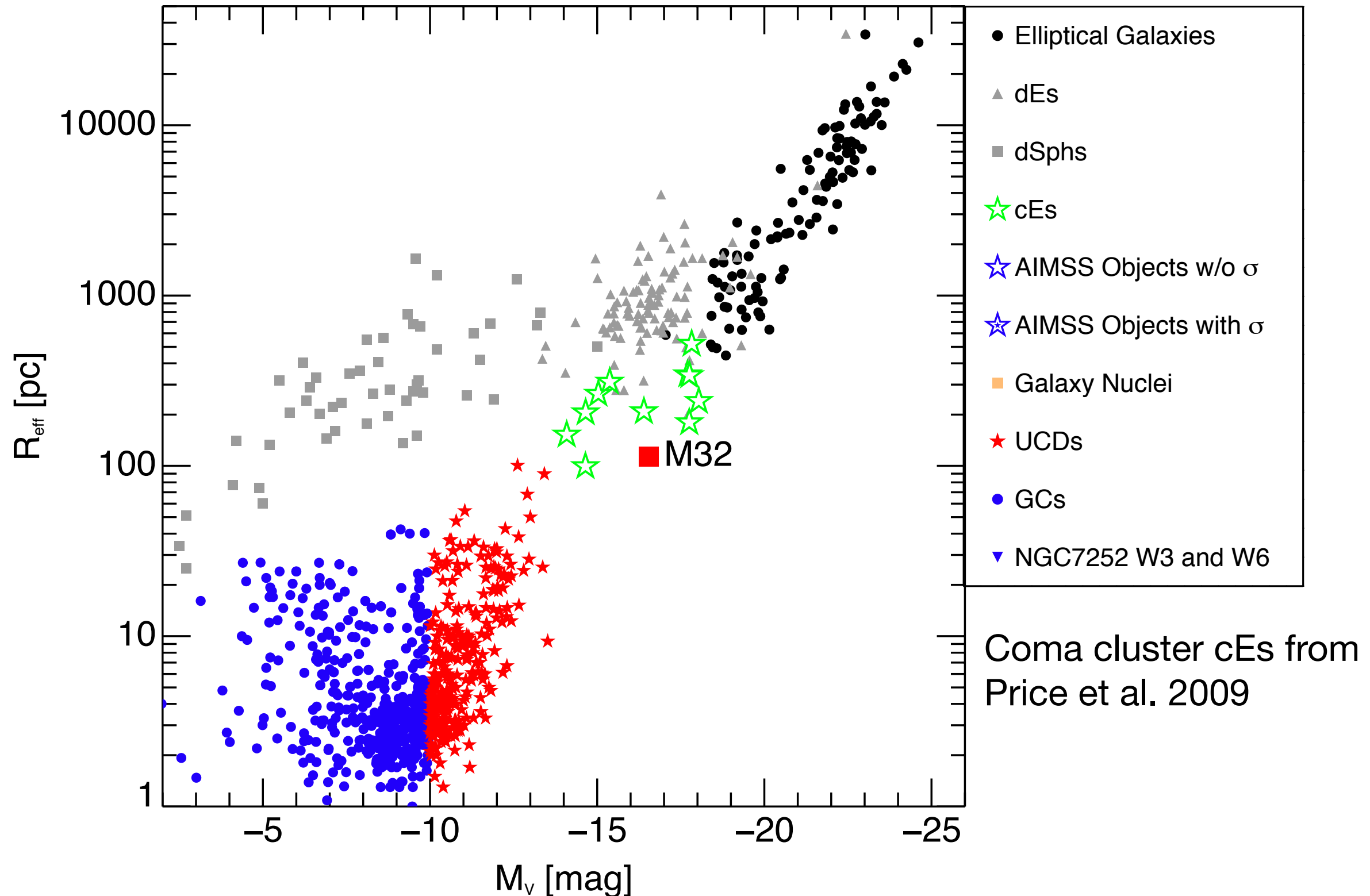
cE Recent History

Plot after Misgeld & Hilker 2011, Brodie et al. 2012

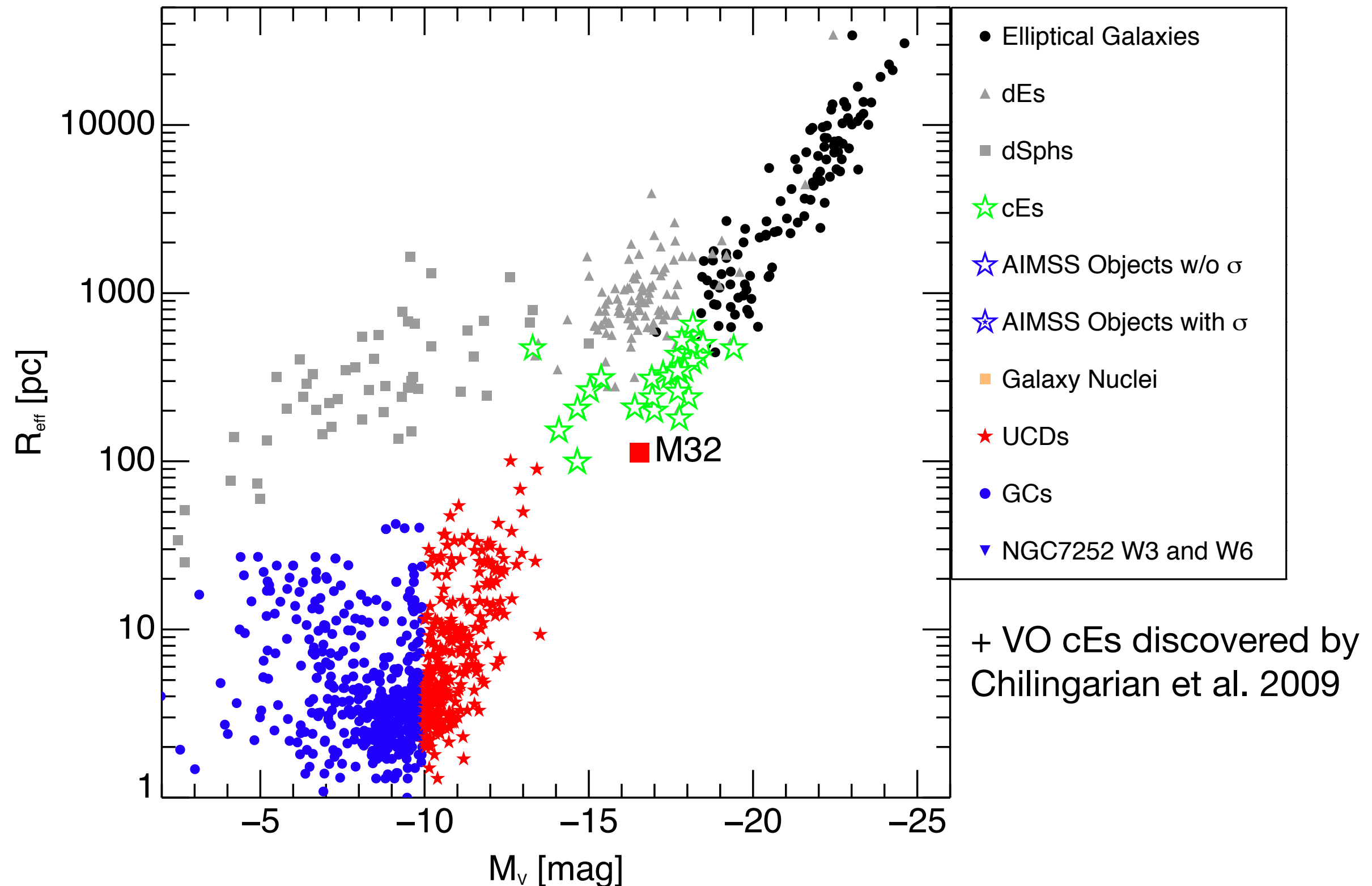


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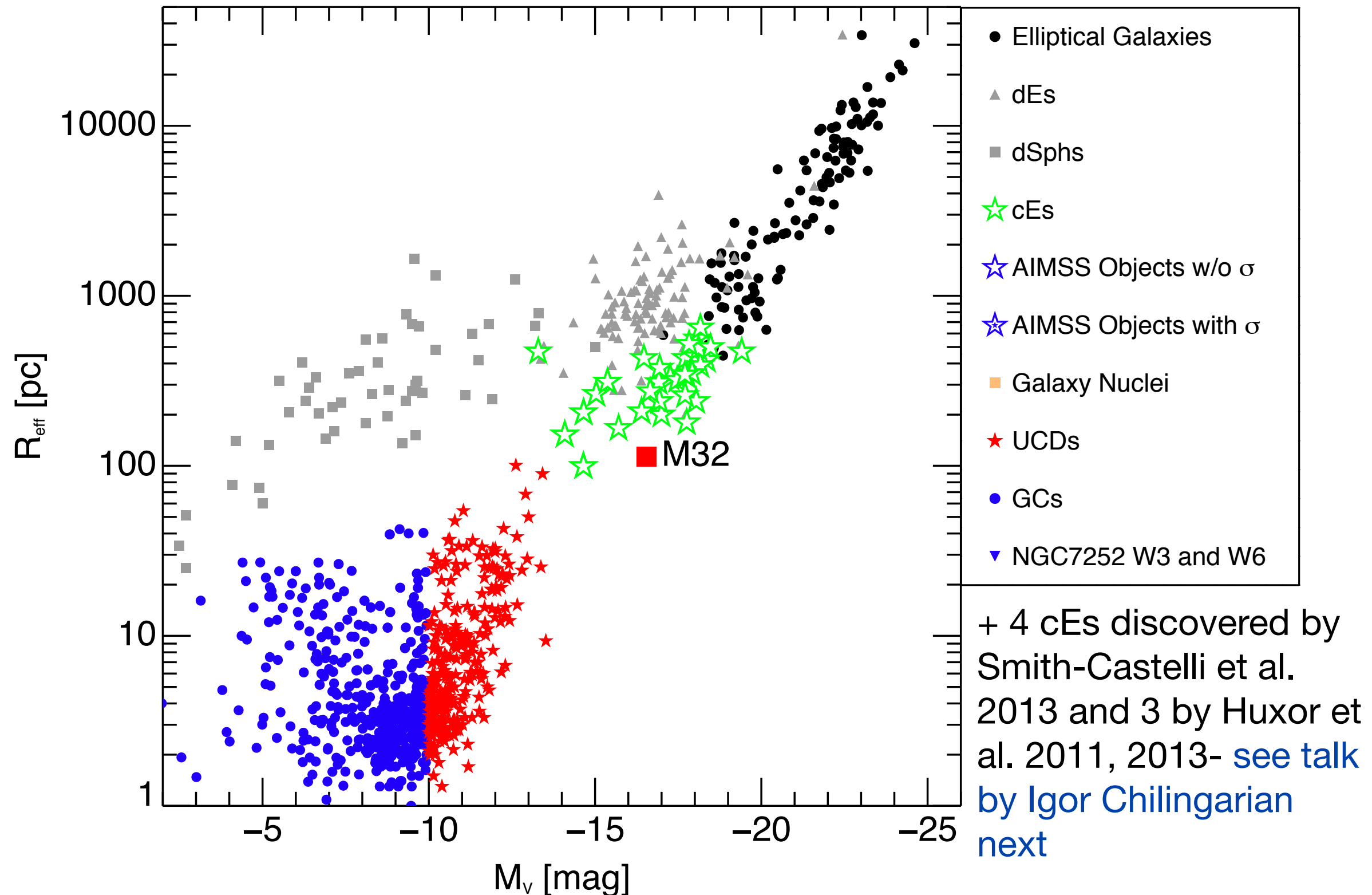
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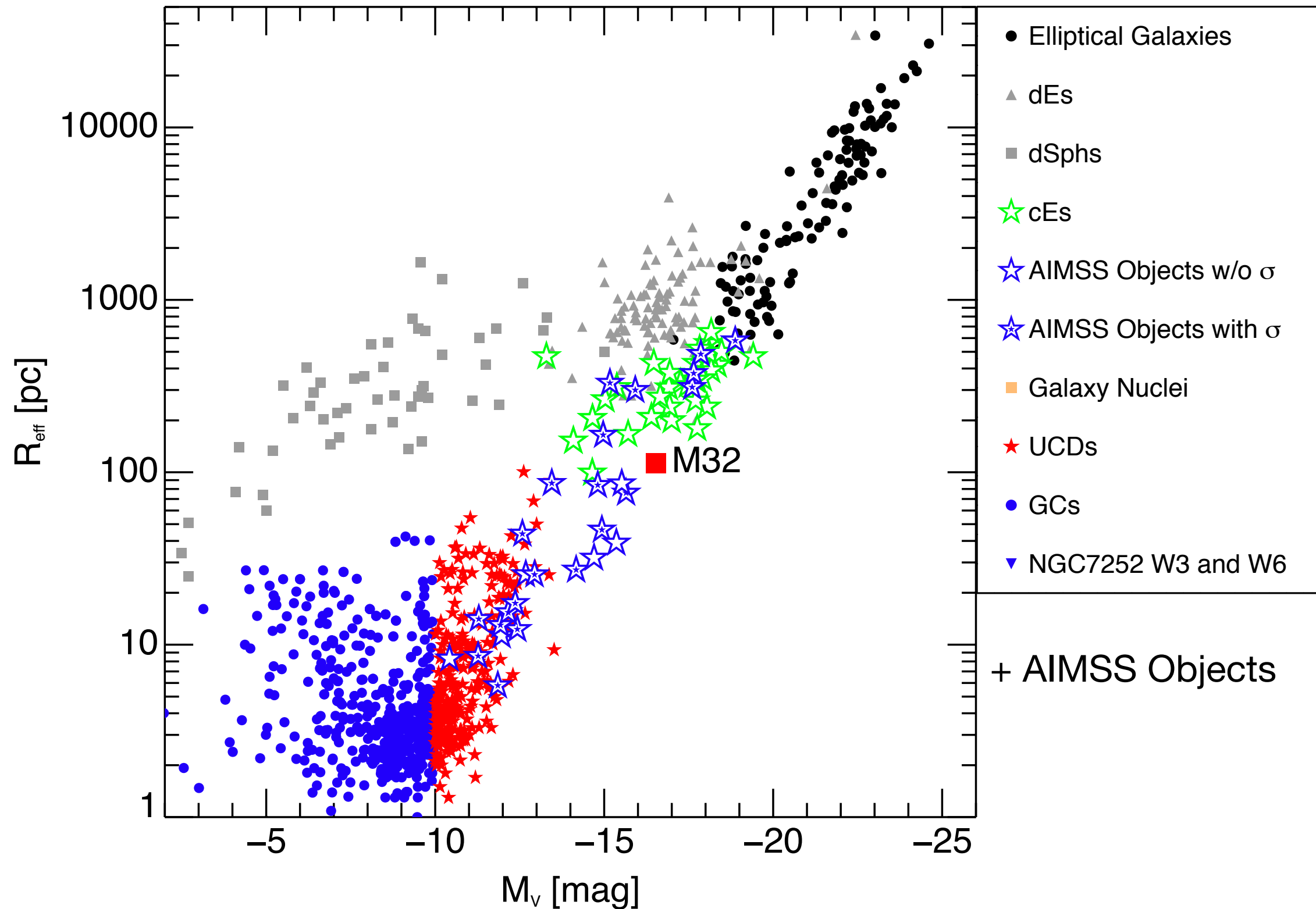
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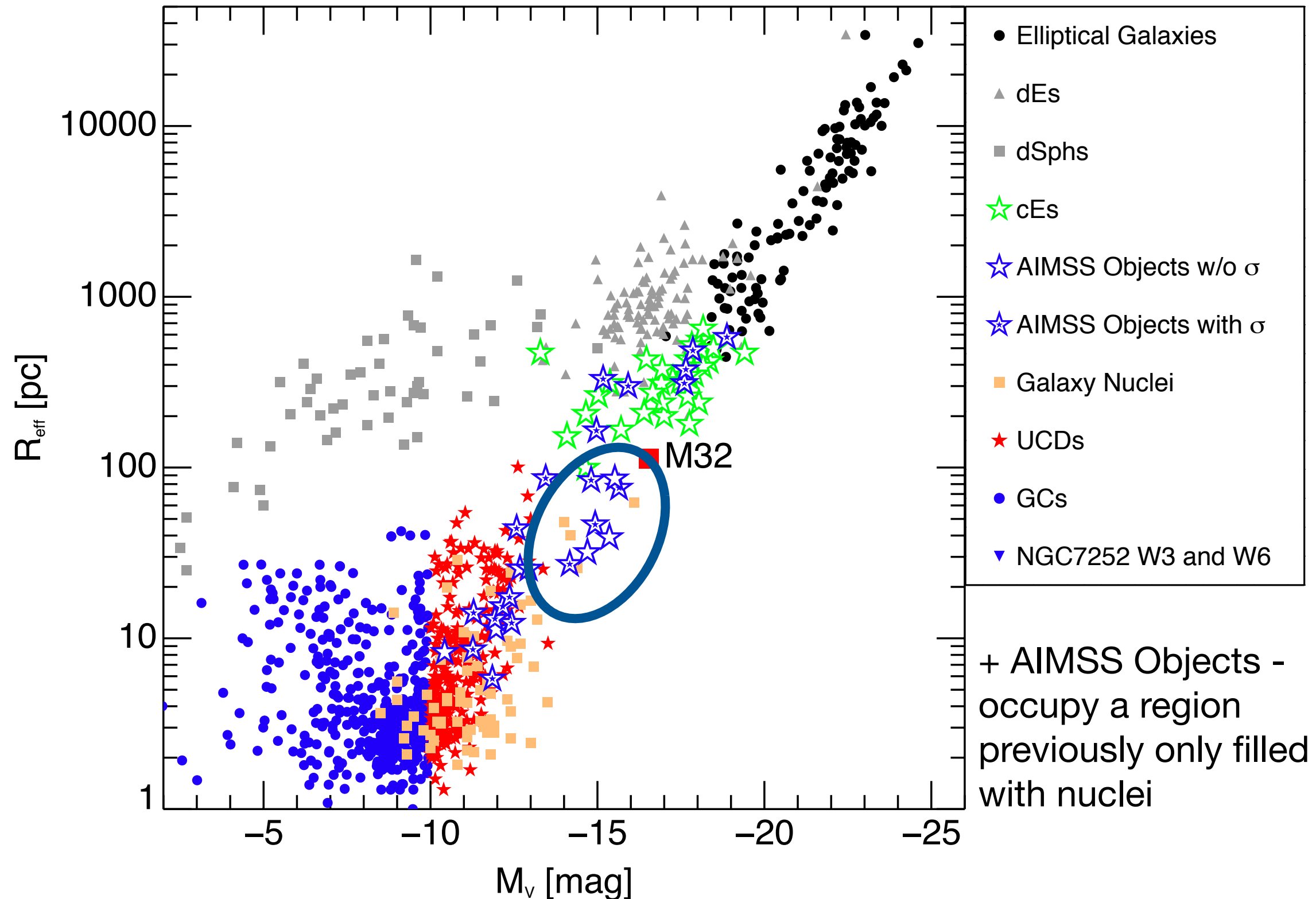
cEs BCT (Before Chilingarians Talk)



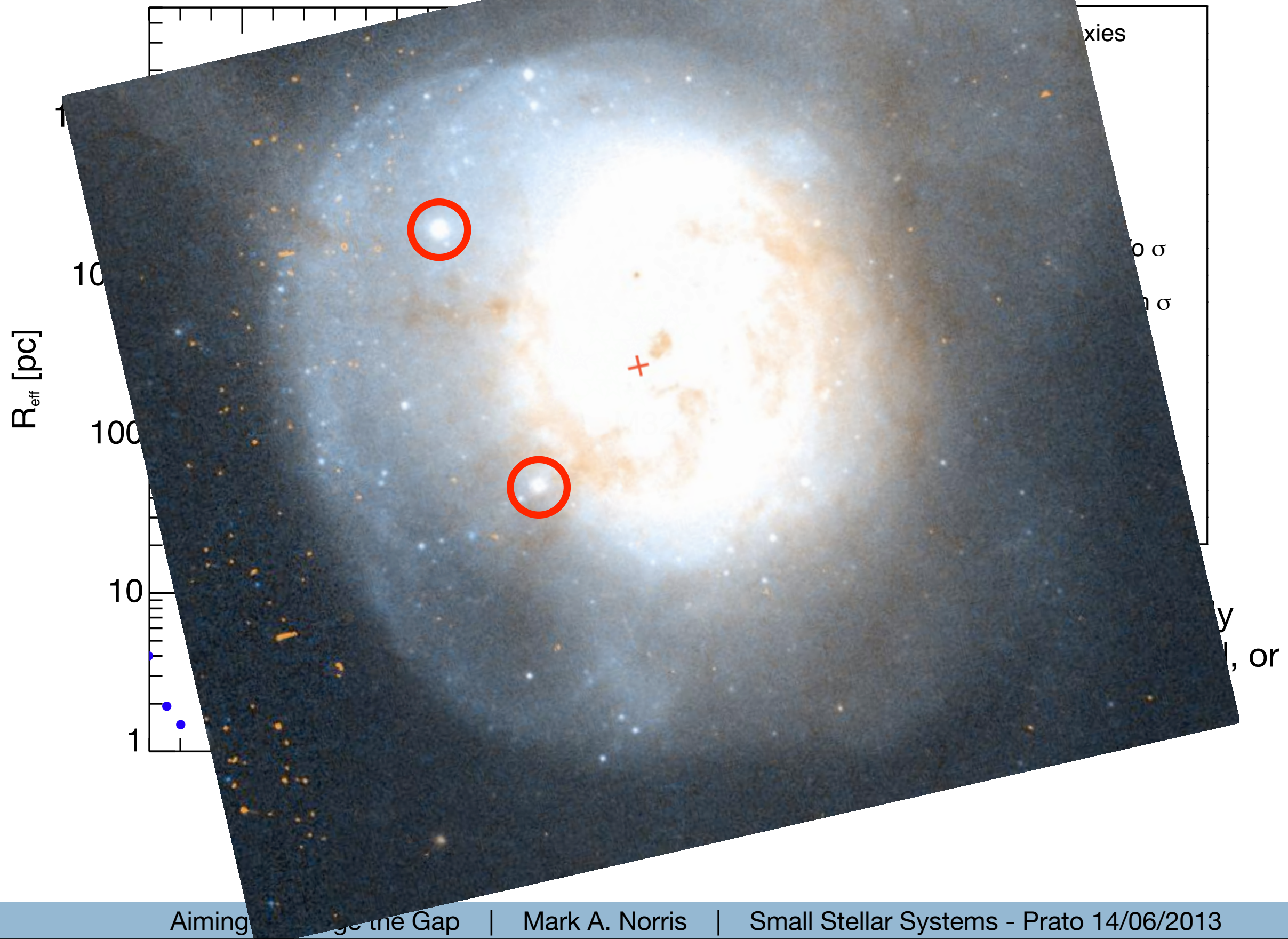
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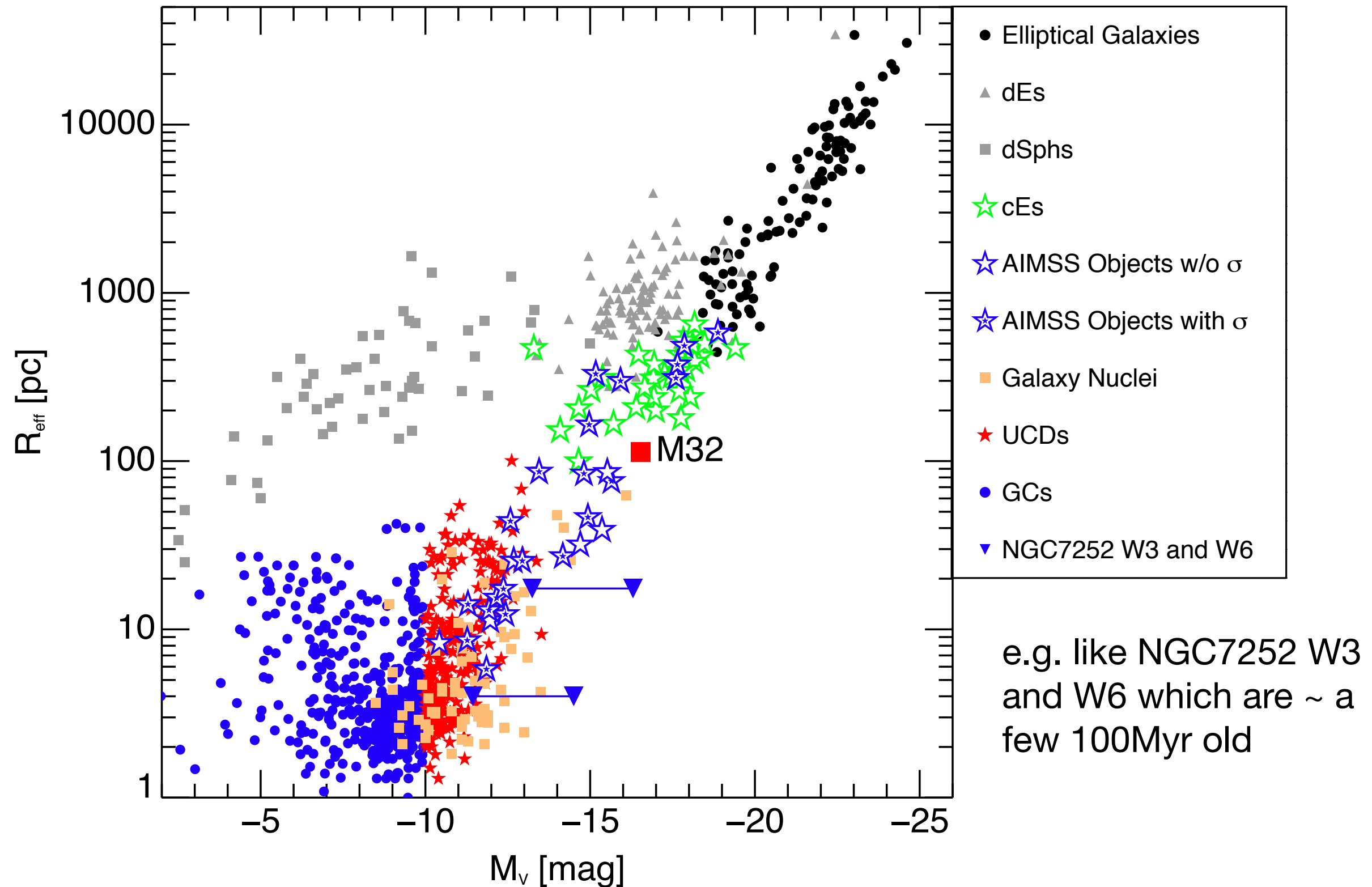
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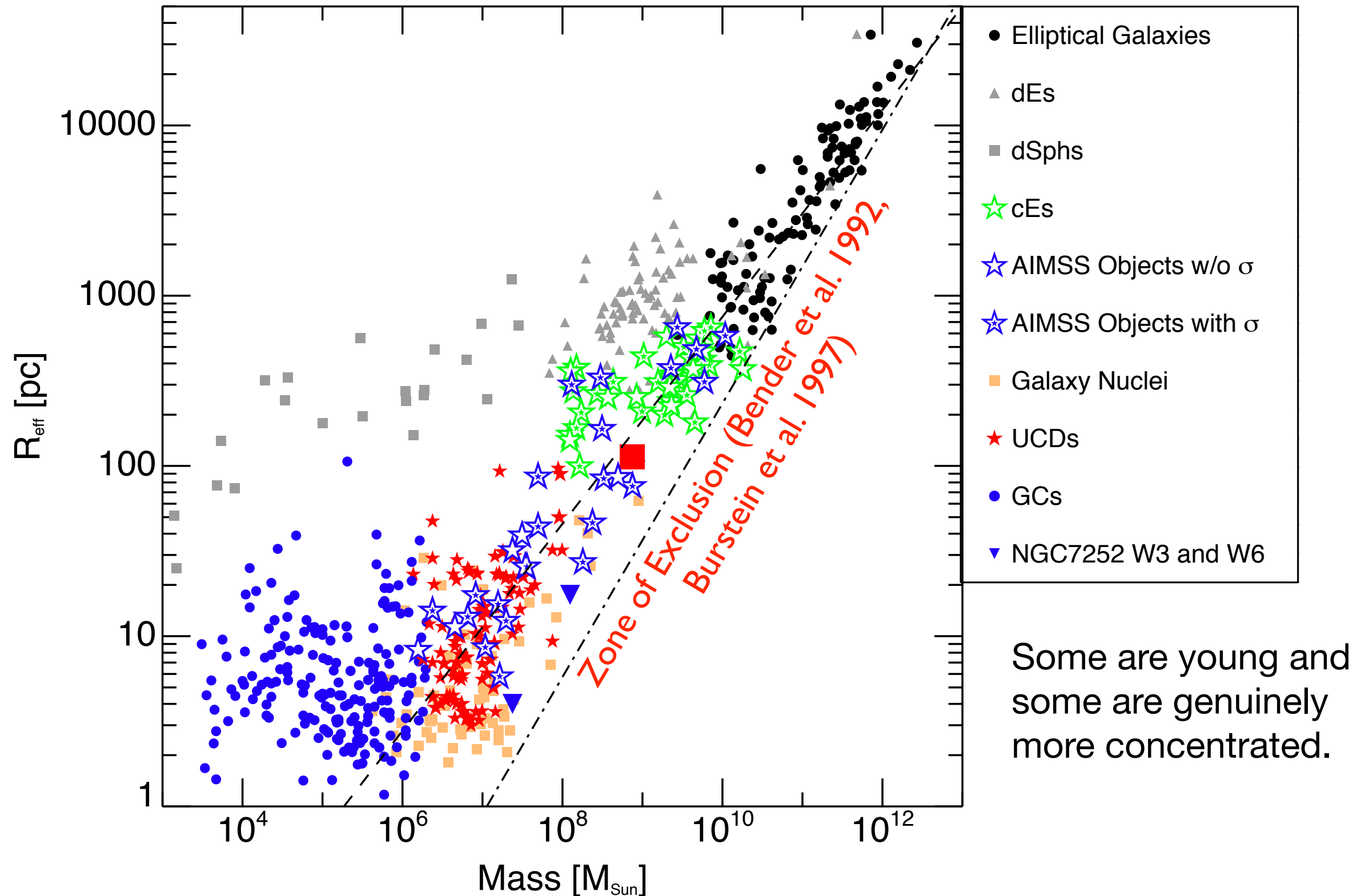
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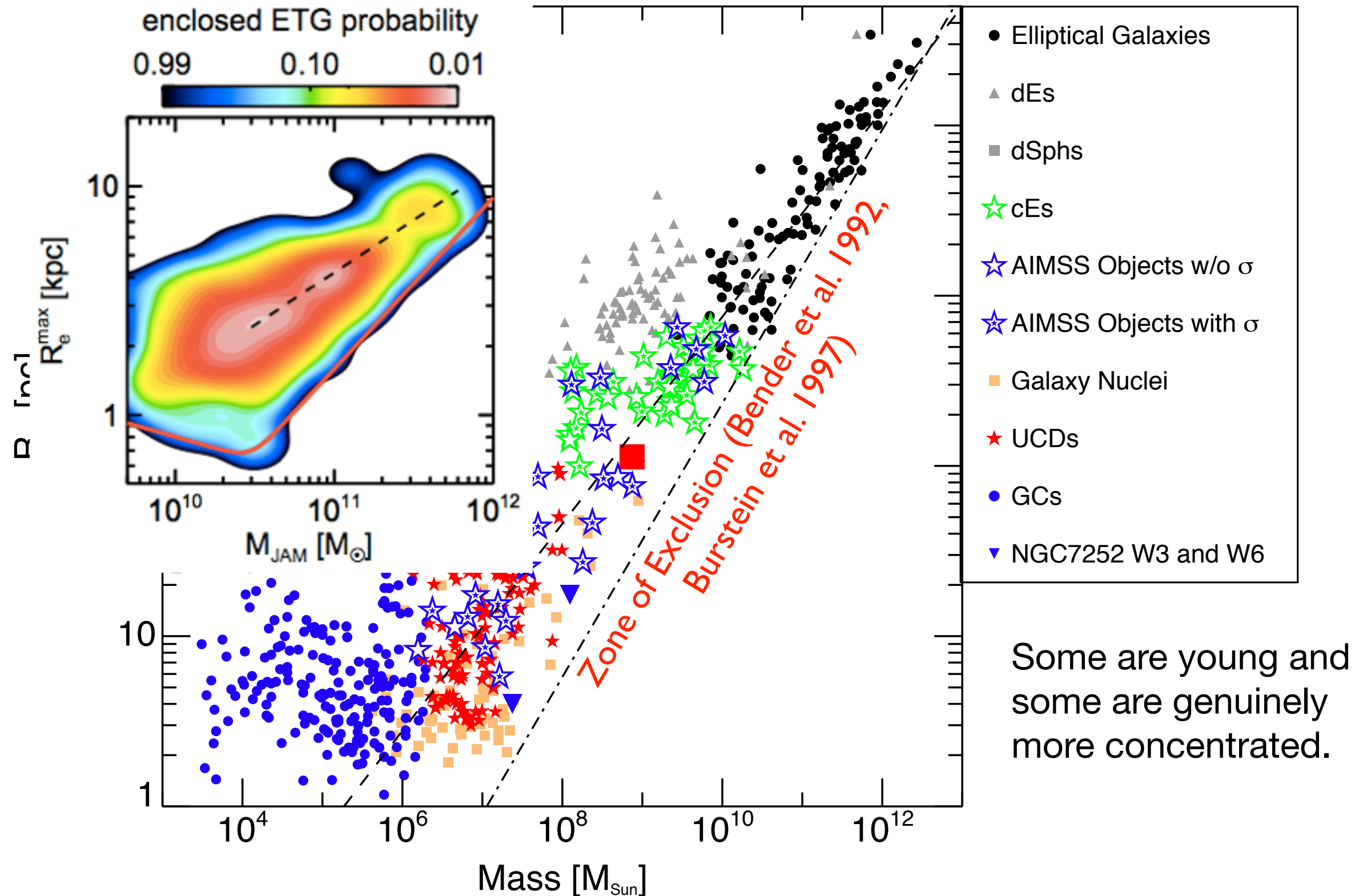


Stellar Mass - Size

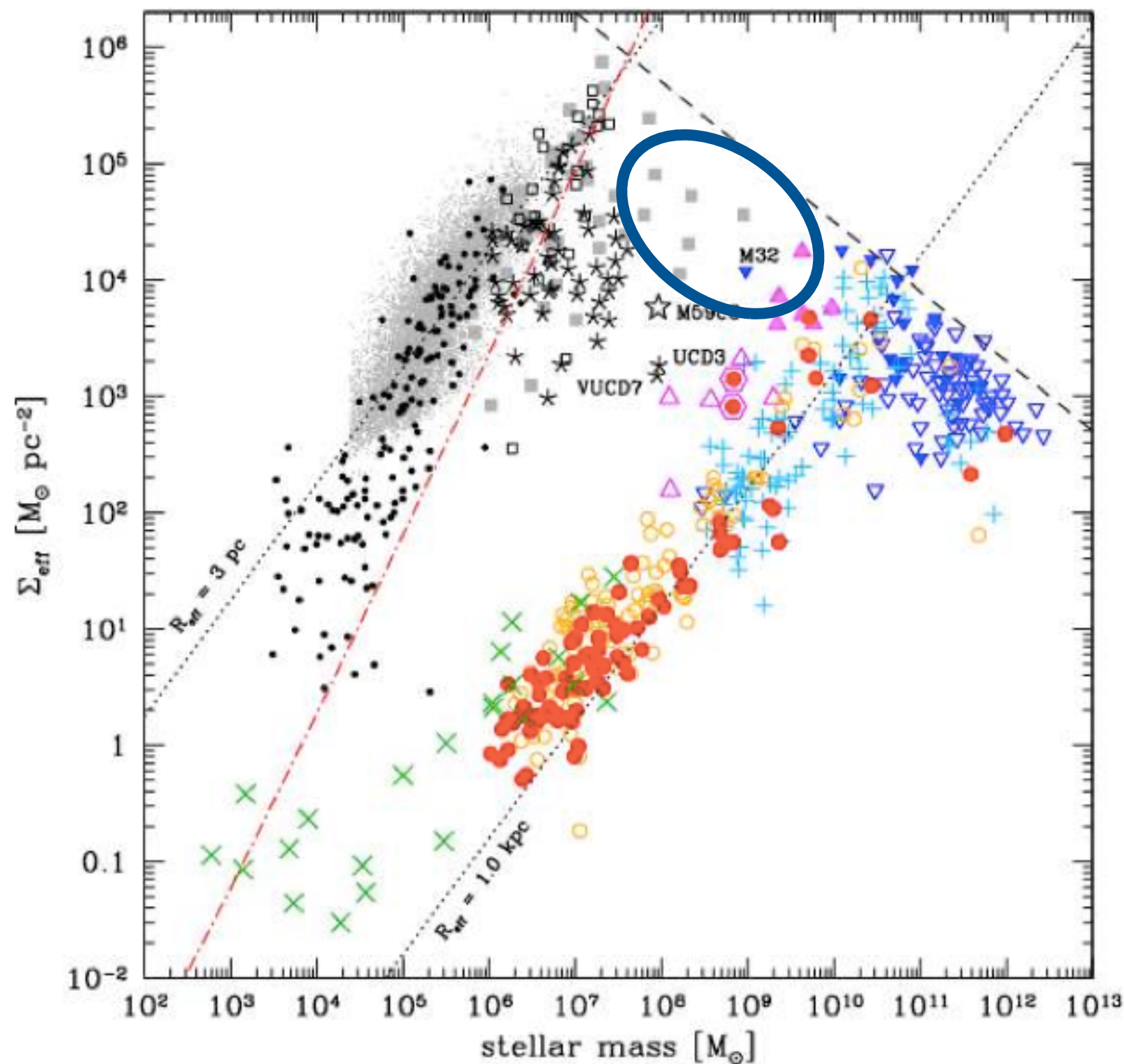


Stellar Mass - Size

Cappellari et al. 2012 (ATLAS^{3D} XX)

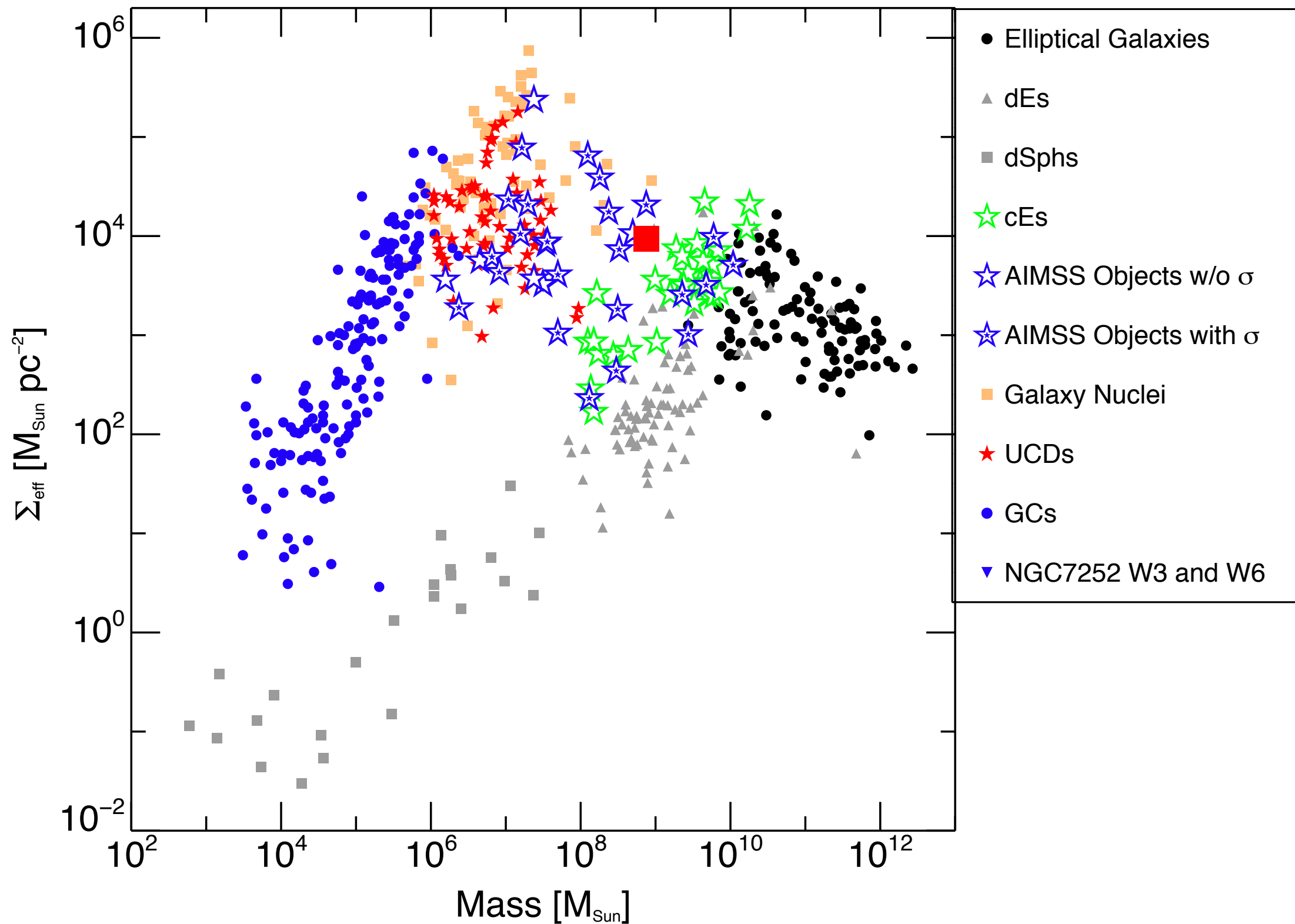


Stellar Mass - Mass Surface Density

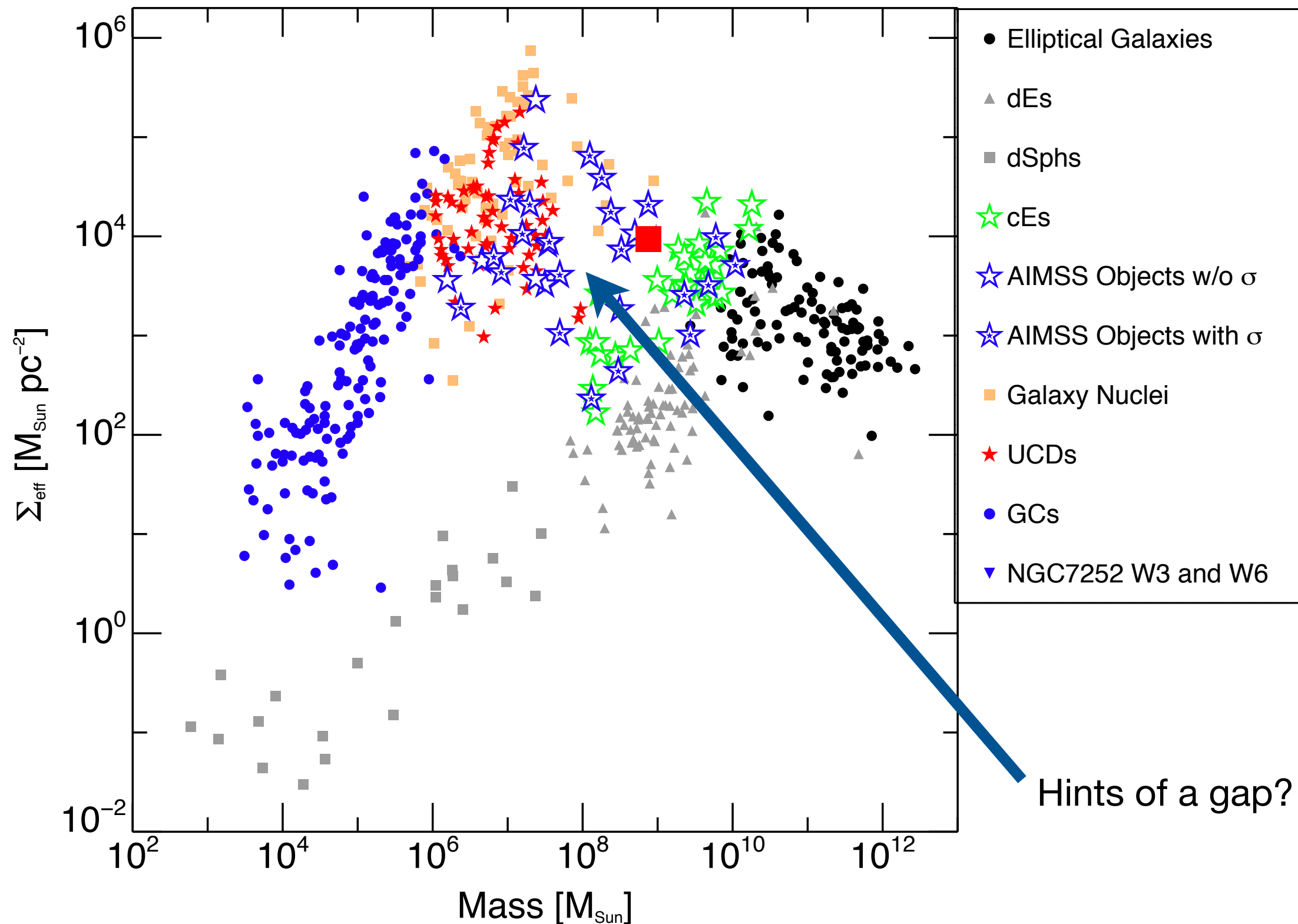


The situation as of
Misgeld & Hilker 2011

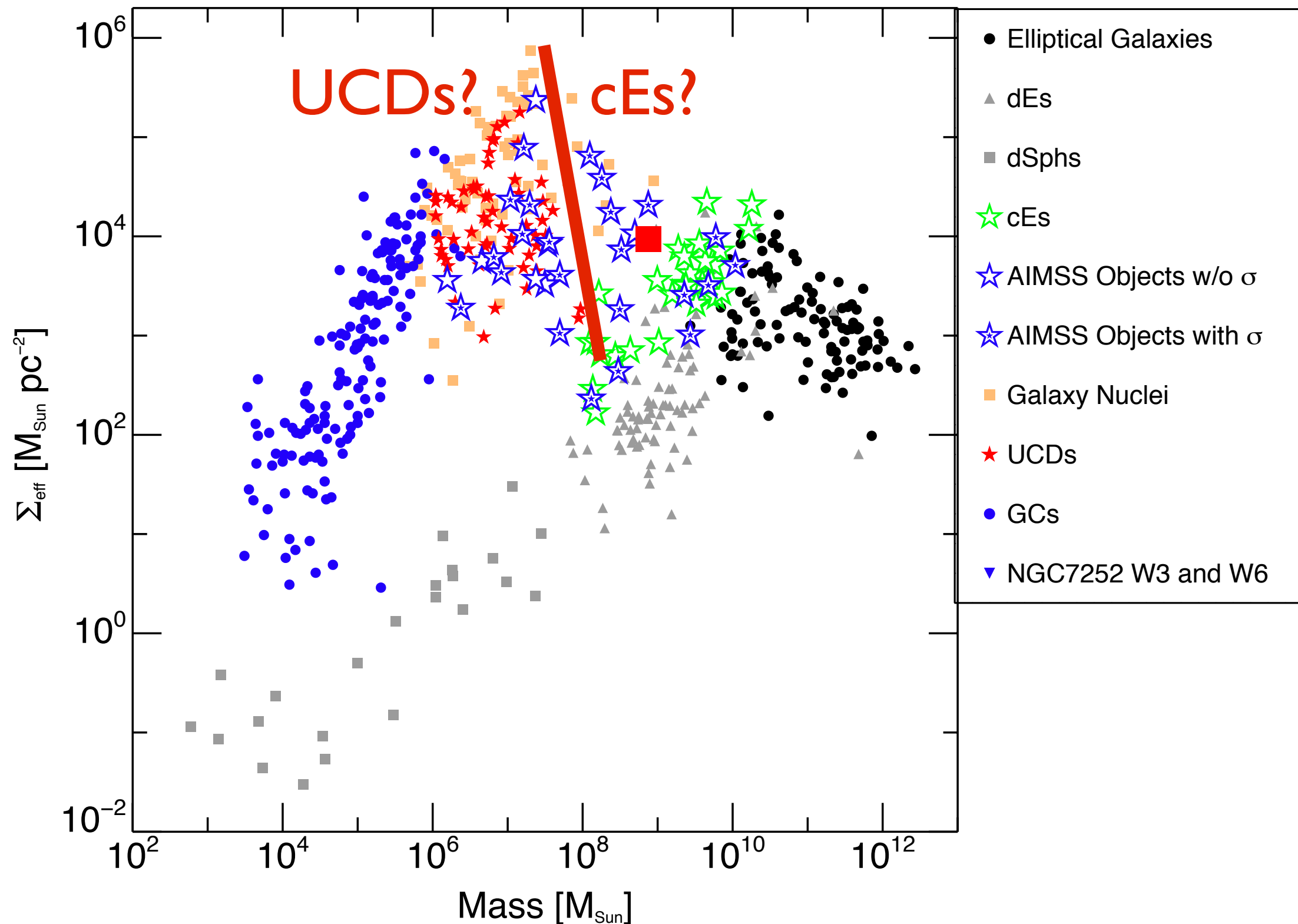
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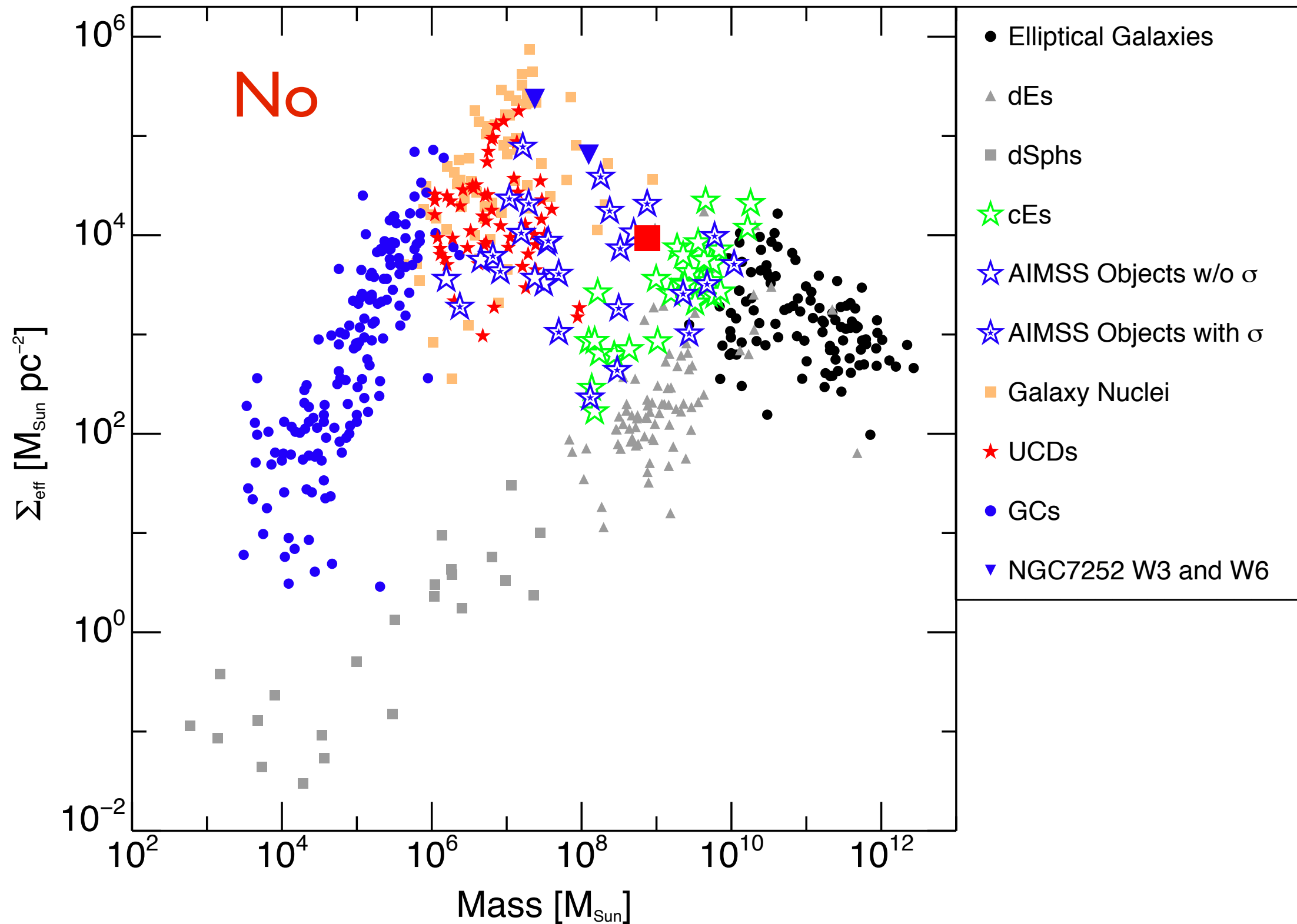
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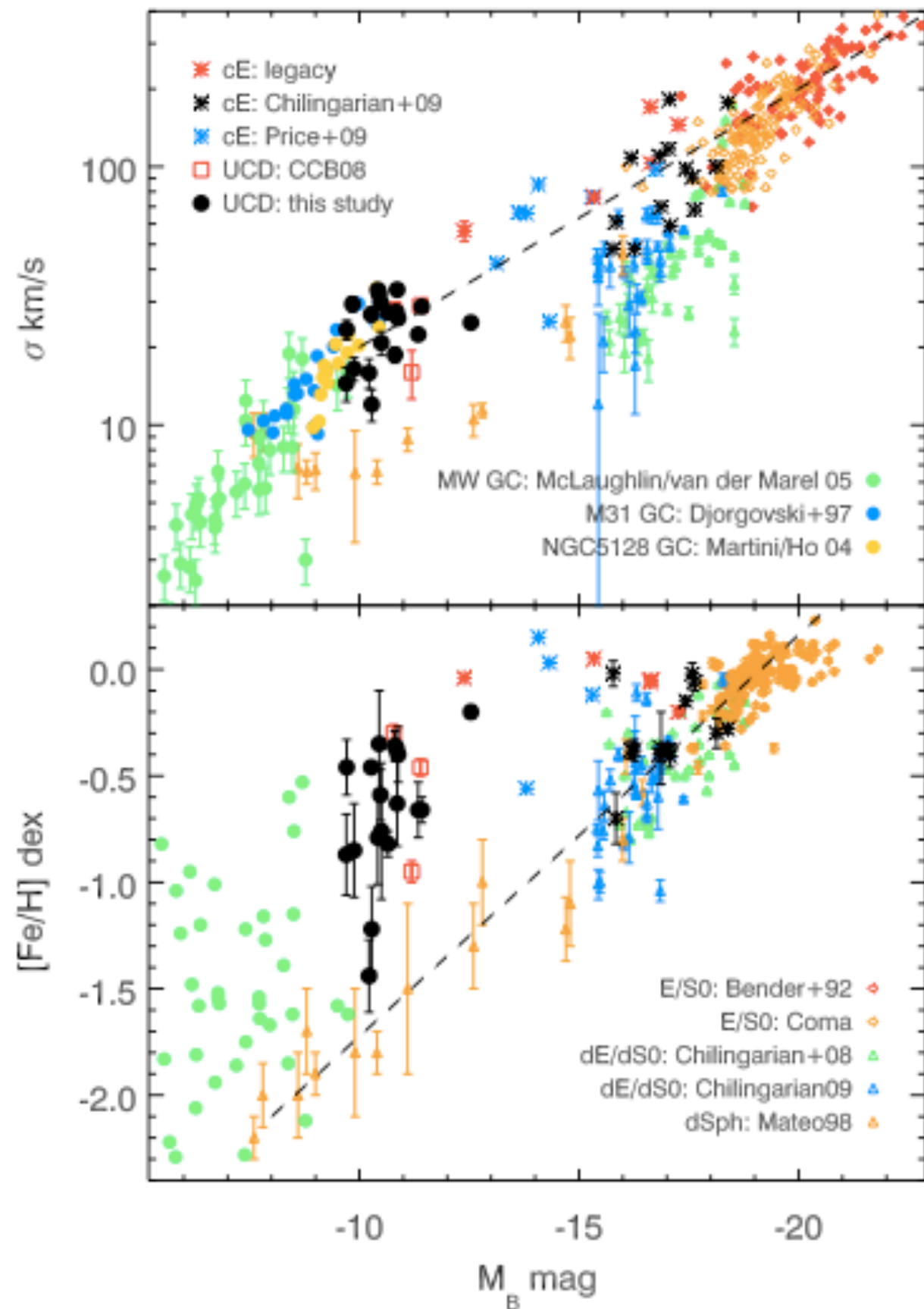
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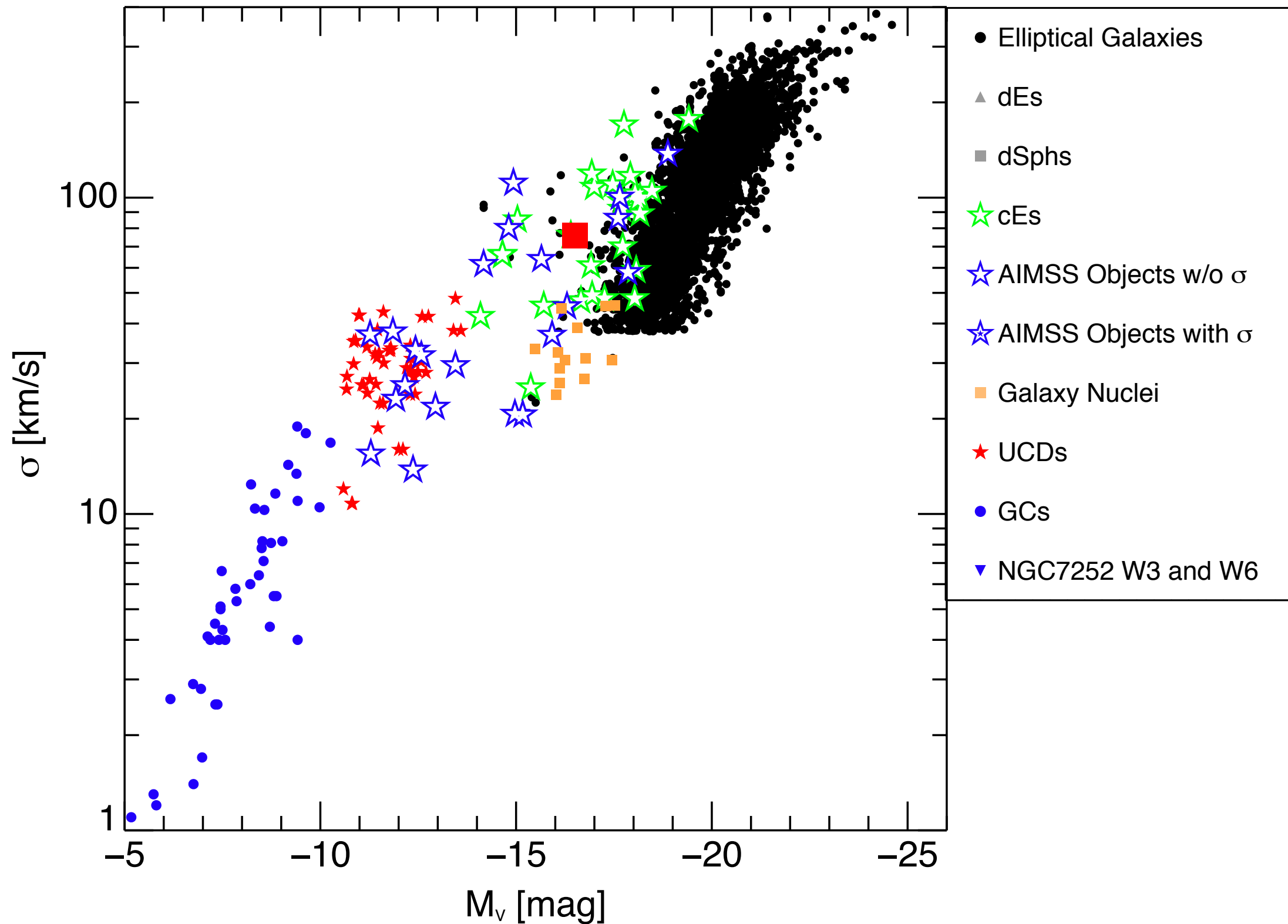
Velocity Dispersion



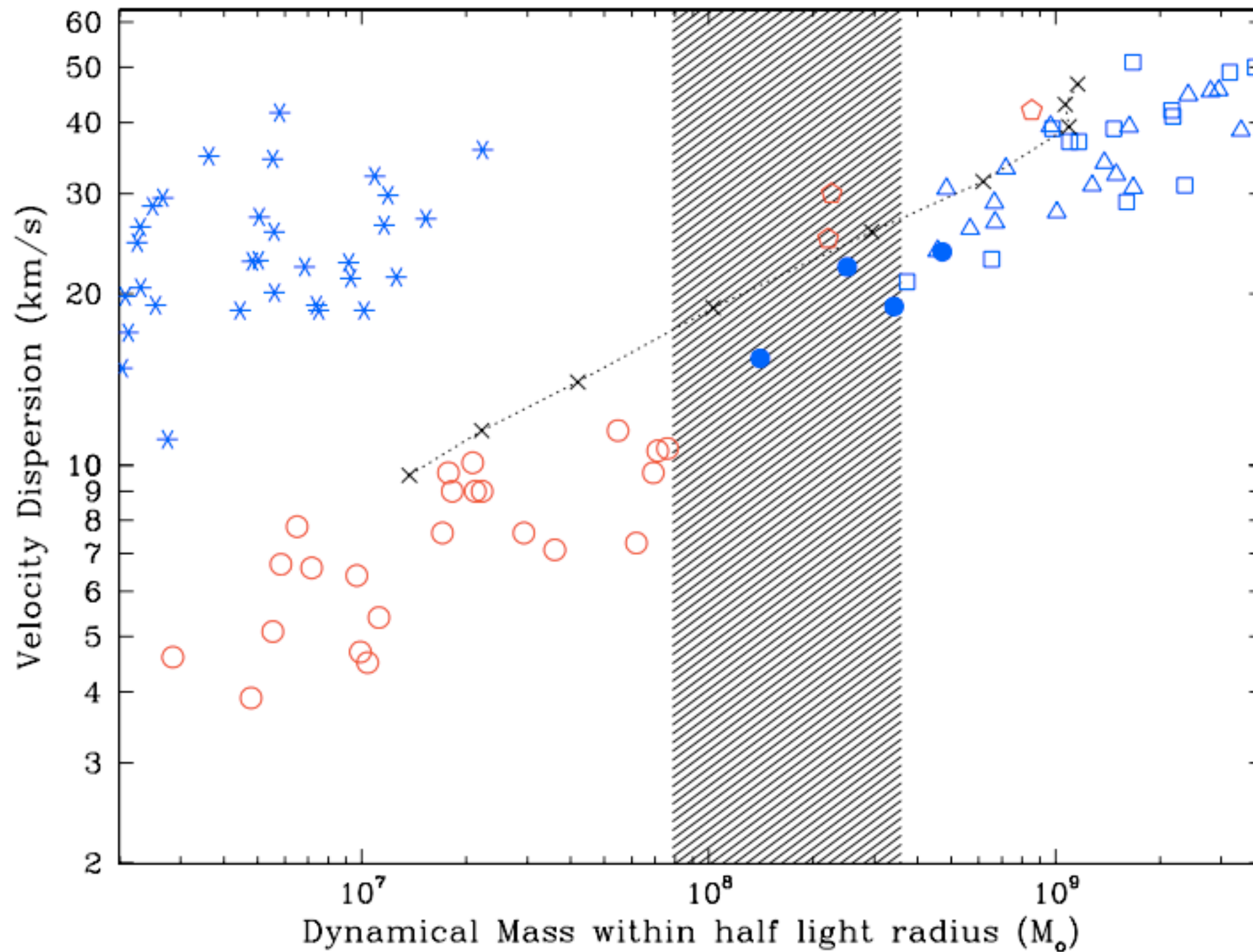
Sigma vs M_B from Chilingarian, Mieske, Hilker, & Infante 2010

Hints of a separation between UCDs/cEs and dEs/dS0s.

Velocity Dispersion



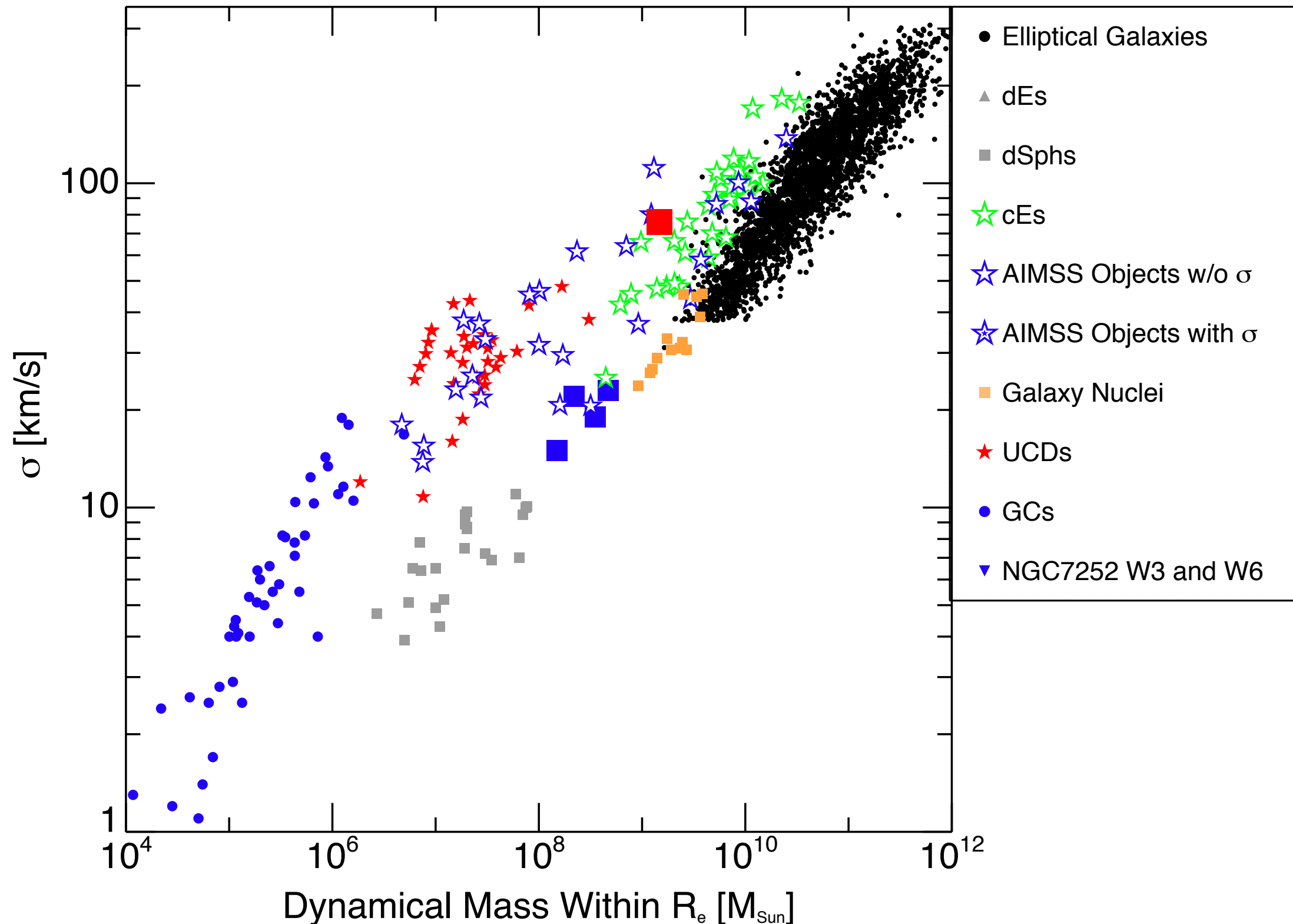
Velocity Dispersion



From Forbes, Spitler, Graham, Foster, Hau & Benson. 2010

Velocity Dispersion

Two sequences in σ - dynamical mass - GCs/UCDs/cEs different from galaxies and nuclei.

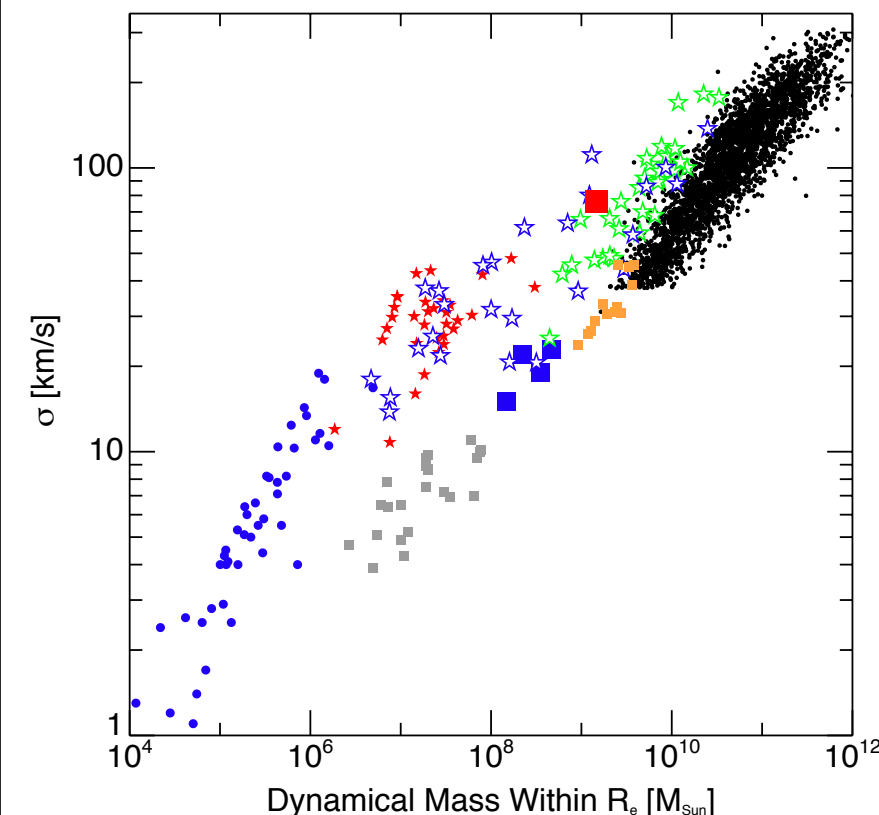
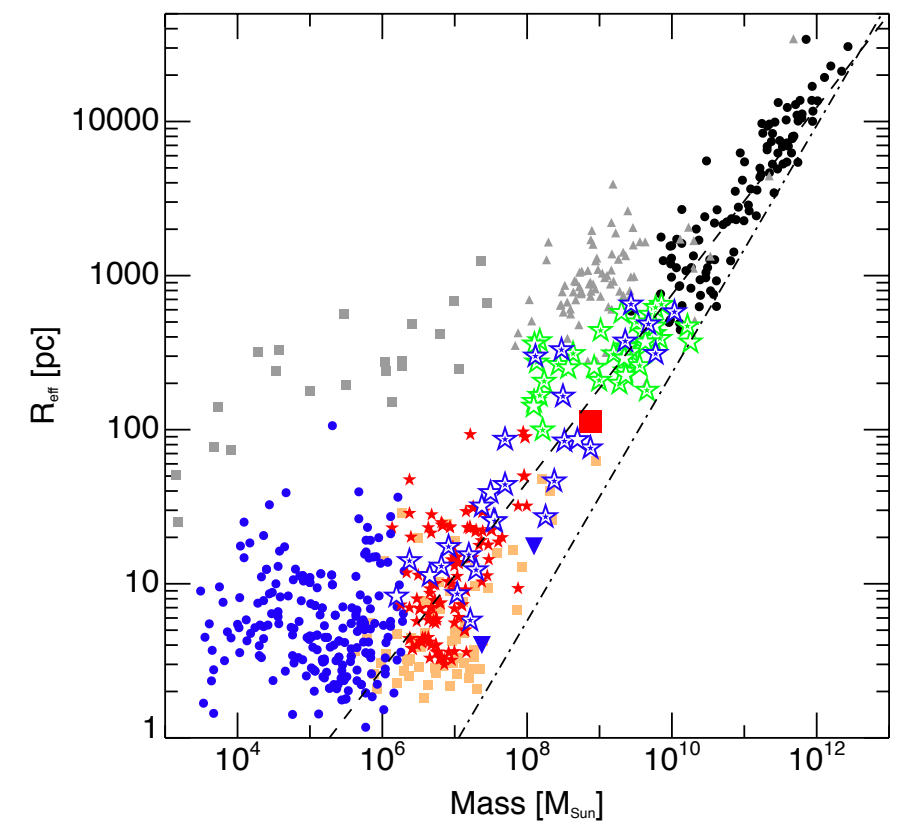


Conclusions

1) We have successfully “bridged the gap” - we now have a continuous sequence of objects from GCs with $R_e = 1\text{pc}$ to gEs with $R_e = 30\text{kpc}$.

2) Our objects are some of the densest free-floating objects known.

Still nothing in the “zone of exclusion”.



3) So far, our cEs seem consistent with a stripping scenario. They display central σ more consistent with objects 10-100 times their dynamical mass.

Ongoing spectroscopy for stellar population information will be able to confirm this conclusion.

Open Questions

- 1) How common are compact ellipticals? - **See next talk by Igor Chilingarian.**
- 2) Can cE's form in more than one way? - **See paper by Huxor, Phillips and Price 2013, for an example of a very isolated cE.**
- 3) Are some cE's related to the massive compact galaxies at intermediate z ?

The AIMSS Catalogue

2) The AIMSS Catalogue - A fully comprehensive catalogue of CSS properties:

- All available photometry, UV to IR
- m - M /distance
- Structural parameters; R_e , Sersic n , King n etc
- Stellar Population parameters, Age, $[Z/H]$, $[\alpha/Fe]$
- V , σ , M_\star , M_{dyn}



WE WANT YOU!R DATA

The AIMSS Catalogue



The AIMSS Catalogue

To help with the legacy value of your data please consider:

- 1) Providing uncertainties! Especially on photometry.**
- 2) Obtaining multi-band photometry in common systems. Ideally SDSS or Johnson-Cousins + 2MASS
- 3) Once published, providing it to me to add to the catalogue, I will be setting up a web form to do this

Thank You